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$Z', \gamma'$

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# Nonexotic $Z'$

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**Nonanomalous  $U(1)_z$  gauge symmetry without new fermions charged under  $SU(3)_C \times SU(2)_W \times U(1)_Y$**

**Allow an arbitrary number of  $\nu_R$ 's**

**Assume:**

- generation-independent charges,
- quark and lepton masses from standard model Yukawa couplings

## Fermion and scalar gauge charges:

	$SU(3)_C$	$SU(2)_W$	$U(1)_Y$	$U(1)_z$
$q_L^i$	3	2	1/3	$z_q$
$u_R^i$	3	1	4/3	$z_u$
$d_R^i$	3	1	-2/3	$2z_q - z_u$
$l_L^i$	1	2	-1	-3 $z_q$
$e_R^i$	1	1	-2	-2 $z_q - z_u$
$\nu_R^k$ , $k = 1, \dots, n$	1	1	0	$z_k$
$H$	1	2	+1	$-z_q + z_u$
$\varphi$	1	1	0	1

$[SU(3)_C]^2 U(1)_z$ ,  $[SU(2)_W]^2 U(1)_z$ ,  $U(1)_Y [U(1)_z]^2$  and  
 $U(1)_Y [U(1)_z]$  anomalies cancel

**Gravitational- $U(1)_z$  and  $[U(1)_z]^3$   
anomaly cancellation conditions:**

$$\frac{1}{3} \sum_{k=1}^n z_k = -4z_q + z_u$$

$$\left( \sum_{k=1}^n z_k \right)^3 = 9 \sum_{k=1}^n z_k^3$$

- For  $n \leq 2$ :

$z_1 = -z_2 \Rightarrow z_u = 4z_q \Rightarrow$  trivial or  **$Y$ -sequential  $U(1)_z$ -charges**

- For  $n \geq 3$ :

$U(1)_{B-L}$  charges:  $z_1 = z_2 = z_3 = -4z_q + z_u$

or  $z_1 = z_2 = -(4/5)z_3 = -16z_q + 4z_u = -4$

$\nu$  masses: three LH Majorana,

two dimension-7 and one dimension-12 Dirac operators,  
RH Majorana ops. of dimension ranging from 4 to 13

or ...

**In general:**  $Z'$  mixes with  $Z$

$\Rightarrow$  LEP I requires  $M_{Z'} \gtrsim 2$  TeV

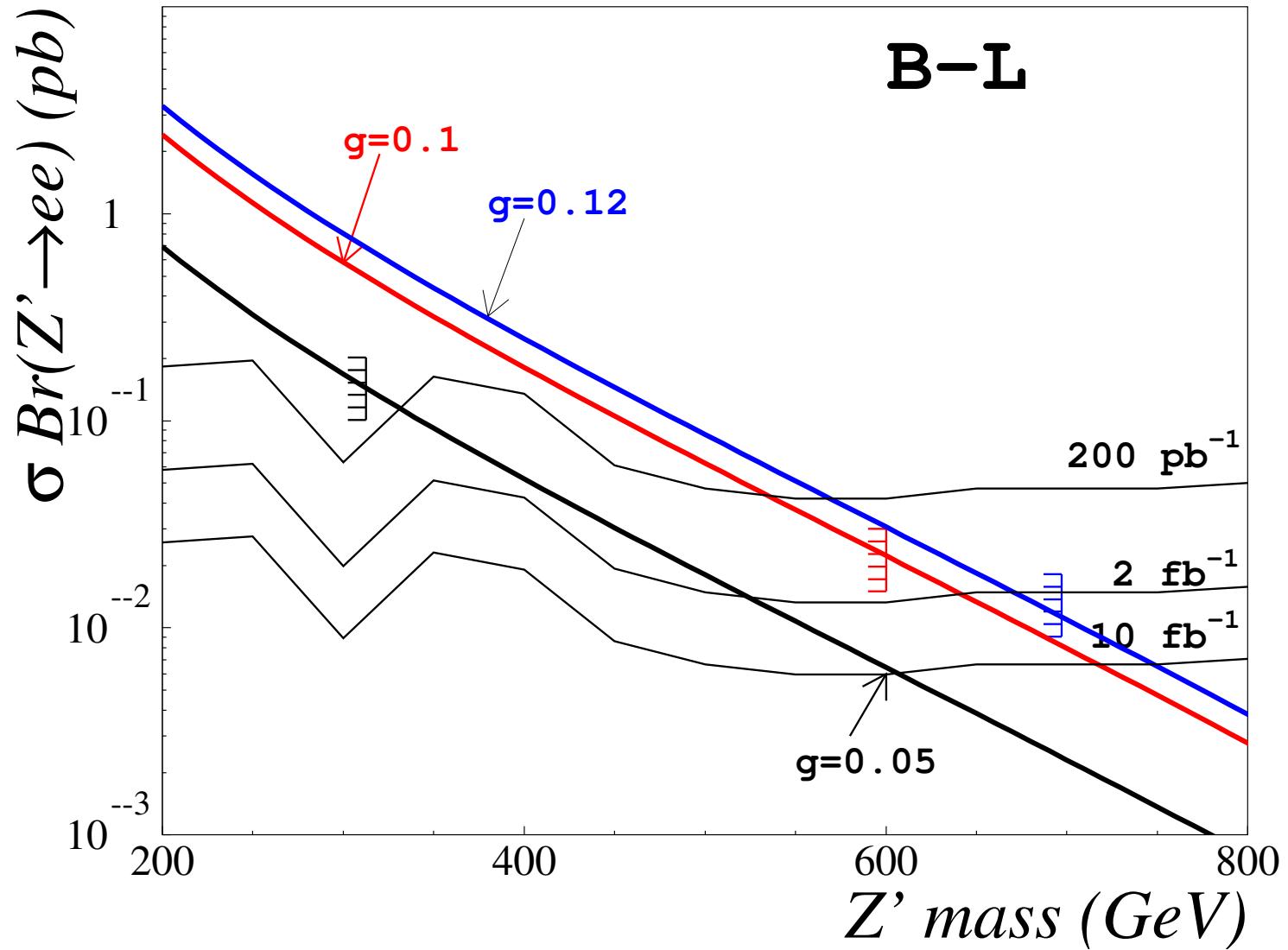
**Special case:**  $SU(3)_C \times SU(2)_W \times U(1)_Y \times \textcolor{red}{U(1)_{B-L}}$

$$z_q = z_u = z_d = -\frac{z_l}{3} = -\frac{z_e}{3} = -\frac{z_\nu}{3} \implies z_H = 0$$

**No  $Z_{B-L}$ - $Z$  mixing at tree level!**

*Best bounds on  $z_l g_z$  come from limits on direct production  
at the Tevatron and at LEP II*

**Z' searches at the Tevatron:**



More general charges are allowed in the presence of new fermions:

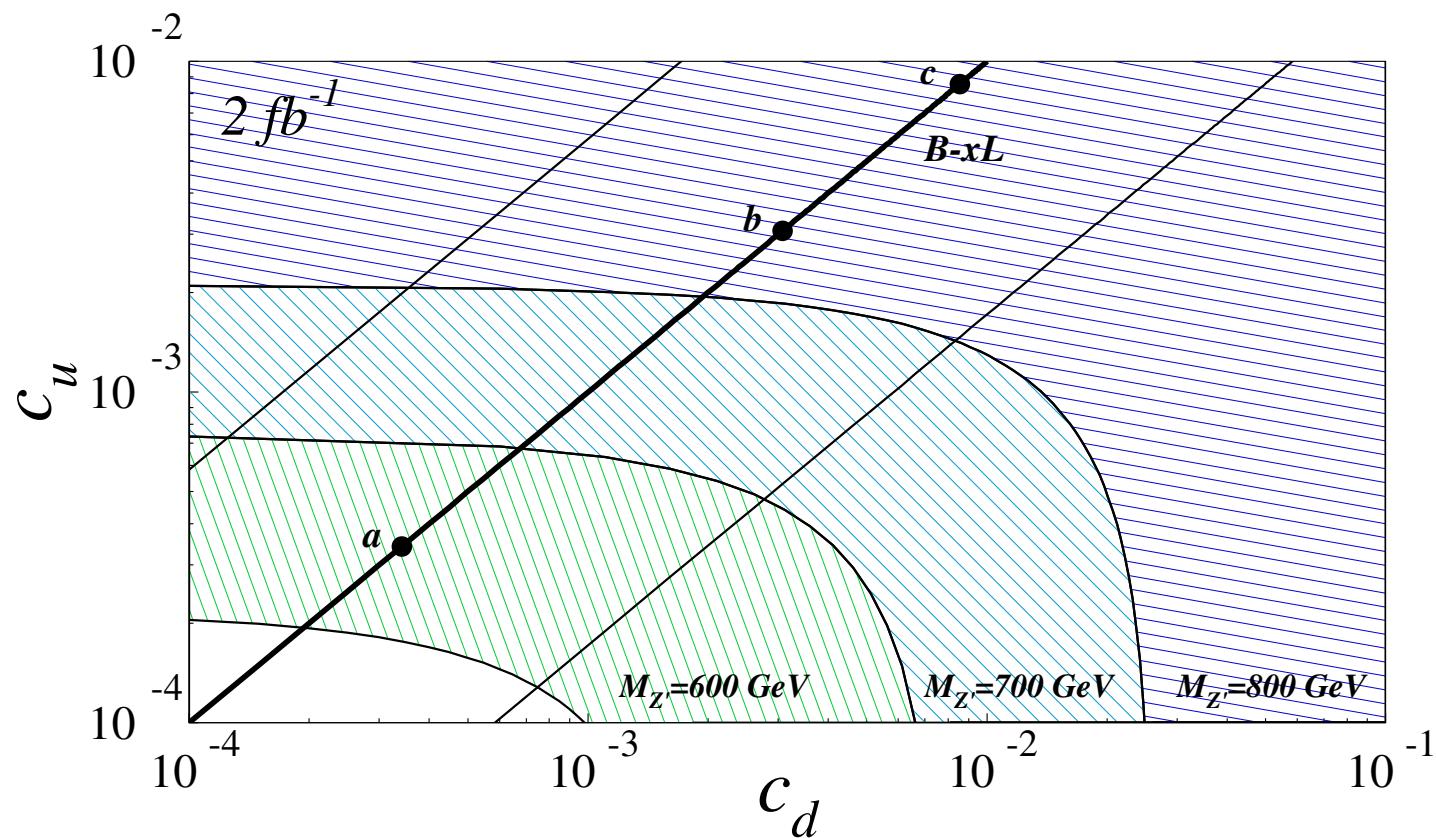
	$SU(3)$	$SU(2)$	$U(1)_Y$	$U(1)_{B-xL}$	$U(1)_{q+xu}$	$U(1)_{10+x\bar{5}}$	$U(1)_{d-xu}$
$q_L$	3	2	1/3	1/3	1/3	1/3	0
$u_R$	3	1	4/3	1/3	$x/3$	-1/3	$-x/3$
$d_R$	3	1	-2/3	1/3	$(2-x)/3$	$-x/3$	1/3
$l_L$	1	2	-1	$-x$	-1	$x/3$	$(-1+x)/3$
$e_R$	1	1	-2	$-x$	$-(2+x)/3$	-1/3	$x/3$
$\nu_R$	1	1	0	-1	$(-4+x)/3$	$(-2+x)/3$	$-x/3$
$\nu'_R$				.	.	$-1-x/3$	.
$\psi_L^l$	1	2	-1	-1	.	$-(1+x)/3$	$-2x/5$
$\psi_R^l$				$-x$	.	2/3	$(-1+x/5)/3$
$\psi_L^e$	1	1	-2	-1	.	.	.
$\psi_R^e$				$-x$	.	.	.
$\psi_L^d$	3	1	-2/3	.	.	-2/3	$(1-4x/5)/3$
$\psi_R^d$				.	.	$(1+x)/3$	$x/15$

## A user-friendly parametrization:

$$\sigma (p\bar{p} \rightarrow Z'X \rightarrow l^+l^-X) = \frac{\pi}{48 s} \left[ c_u w_u \left( \frac{M_{Z'}^2}{s}, M_{Z'} \right) + c_d w_d \left( \frac{M_{Z'}^2}{s}, M_{Z'} \right) \right]$$

All the information about charges is contained in:

$$c_{u,d} = g_z^2 (z_q^2 + z_{u,d}^2) \text{Br}(Z' \rightarrow l^+l^-)$$



**Could there exist  
massless gauge bosons  
other than the photon?**

$U(1)_{B-L}$  is the only global symmetry of the standard model that can be gauged and unbroken.

$Z_{B-L}$  coupling to ordinary matter:  $N_n g_z$   
( $N_n$  = number of neutrons)

To avoid deviations from Newton's law:

$$g_z \ll \frac{m_n}{M_{\text{Pl}}} \sim 10^{-19}$$

Tests of the equivalence principle:  $g_z < 10^{-24}$   
weaker limit if  $B - L$  force is screened by  $\nu$ 's (Okun, et al)

*But even when  $z_q = z_u = z_d = z_l = z_e = 0$*

**there can still be interactions of the standard model fields with the new massless gauge boson:**

**higher-dimensional operators!**

$\gamma'$  couplings to leptons:

$$\frac{1}{M^2} P_{\mu\nu} (\bar{l}_L \sigma^{\mu\nu} C_e H e_R + \text{h.c.})$$

$C_e$ : **3  $\times$  3 matrix in flavor space,  
dimensionless parameters**

$\gamma'$  couplings to quarks: similar dimension-6 operators

**In the mass eigenstate basis**  $C_e \rightarrow C'_e = U_L^e C_e U_R^{e\dagger}$

$U_L^e$  and  $U_R^e$  are the unitary matrices that diagonalize the masses of the electrically-charged leptons.

- **Interactions of the mass-eigenstate leptons with  $P^\mu$  (chirality-flip operators  $\sim v_h \approx 174$  GeV) :**

$$\frac{v_h}{M^2} P_{\mu\nu} \bar{e}' \sigma^{\mu\nu} (\text{Re}C'_e + i \text{Im}C'_e \gamma_5) e'$$

→ magnetic-like and electric-like dipole moments

$\text{Re}(C'_e)^{ij}$ ,  $\text{Im}(C'_e)^{ij}$  could have any value  $\lesssim 4\pi$ , but:

chirality-flip operators  $\rightarrow$  probably  $|(C_e)_{ij}| \lesssim |\lambda_e^{ij}|$

$$\implies |(C'_e)_{ij}| \lesssim \frac{m_\tau}{v_h} \approx 10^{-2}$$

$|(C'_e)_{11}|$  may naturally be below  $m_e/v_h \approx 3 \times 10^{-6}$

**Kinetic mixing of  $U(1)_Y \times U(1)_z$  gauge bosons:**

$c_0 B^{\mu\nu} P_{\mu\nu}$     dimension-four operator!

*Holdom 1985:*

**Kinetic terms can be diagonalized and canonically normalized by a  $SL(2, R)$  transformation.**

**Global  $SO(2)$  symmetry:** linear combination of  $U(1)$  fields that couples to hypercharge is the real  $B^\mu$ .

**Orthogonal combination (“paraphoton” =  $\gamma'$ ) does not have any renormalizable couplings to standard model fields.**

*Conclusion:*

*kinetic mixing has no effect on the standard model fields other than a renormalization of the hypercharge gauge coupling.*

## Bosonic interactions of the paraphoton:

$$\frac{1}{M^2} H^\dagger H (c_1 B_{\mu\nu} + \tilde{c}_1 \tilde{B}_{\mu\nu} + c_2 P_{\mu\nu} + \tilde{c}_2 \tilde{P}_{\mu\nu}) P^{\mu\nu}$$

- renormalize the  $U(1)$  gauge couplings
- include vertices with two  $U(1)$  gauge bosons and Higgs bosons.

*These are all operators of  $d \leq 6$  involving both  
 $\gamma'$  and Standard Model fields*

**The strength of the  $\gamma'$  interaction with the electrons depends on**

$$c_e \equiv \frac{v_h}{m_e} |(C'_e)_{11}| \lesssim O(1)$$

**Similar parameters defined for interactions such as  $\mu^+\mu^-\gamma'$ ,  $\mu^\pm e^\mp\gamma'$ , ...**

**Various measurements set limits on these parameters.**

# Primordial Nucleosynthesis

*Constraints on new particles with mass below several MeV.*

**Maximum number of new relativistic degrees of freedom:**

$$\Delta g_*^{\max} = \frac{7}{8} \Delta N^{\max} ; \quad \text{at the } 2\sigma \text{ level: } \Delta N_{\nu}^{\max} \approx 0.6$$

**$\gamma'$  must go out of equilibrium at  $T_P > T_{\text{BBN}} \approx 1 \text{ MeV}$**

**Number of degrees of freedom contributed by  $\gamma'$  during BBN:**

$$\Delta g_*(T_{\text{BBN}}) = 2 \left[ \frac{g_*(T_{\text{BBN}})}{g_*(T_P)} \right]^{4/3} \rightarrow g_*(T_P) > \frac{20.0}{(\Delta N_\nu^{\text{max}})^{3/4}} \gtrsim 30$$

**Freeze-out temperature:**

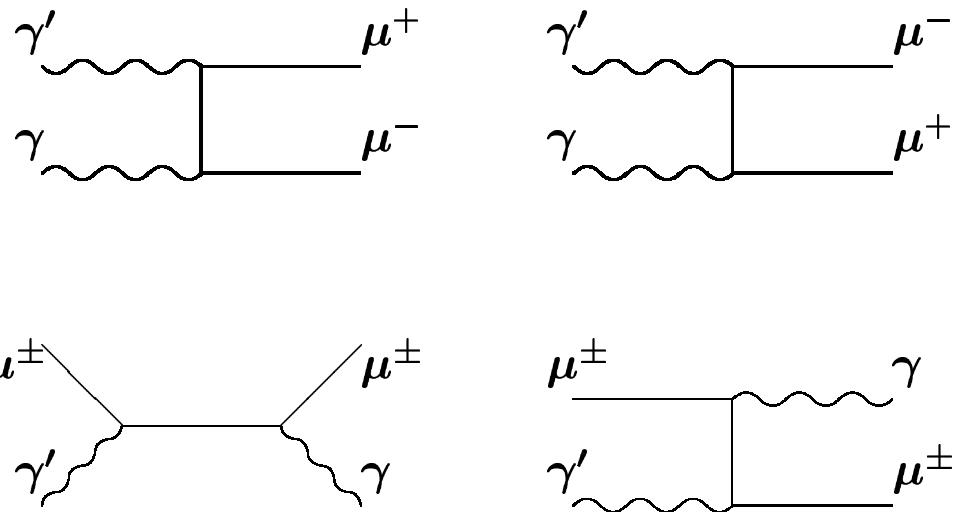
$$T_P > T_{\text{QCD}} \approx 150 - 180 \text{ MeV}$$

$$g_*(T_P) = 247/4$$

(for  $T_P \lesssim T_{\text{QCD}}$ :  $g_* = 69/4$  is too small)

At  $T \approx 200$  MeV:  $\gamma', \gamma, e, \mu, u, d, g, \nu_e, \nu_\mu, \nu_\tau$  are in equilibrium.

Dominant  $\gamma'$  annihilation channels:



Interaction rate of  $\gamma'$ :  $\Gamma = n_{\gamma'} \langle \sigma |v| \rangle$

Number density of  $\gamma'$ :  $n_{\gamma'} = \frac{2\zeta(3)}{\pi^2} T^3 \approx 0.24 T^3$

Thermally averaged cross section:  $\langle \sigma |v| \rangle \sim \frac{\alpha c_\mu^2 m_\mu^2}{M^4}$

**Expansion rate of the universe:**  $H \approx \frac{T^2}{M_{\text{Pl}}} \left( \frac{2\pi^3}{45} g_*(T) \right)^{1/2}$

**At the freeze-out temperature:**  $\Gamma_s \approx H$

$$\Rightarrow \frac{M}{\sqrt{c_\mu}} \approx 3.9 \text{ TeV} \times [g_*(T_P)]^{-1/8} \left( \frac{T_P}{1 \text{ GeV}} \right)^{1/4}$$

**Measurements of light element abundances set a limit on the effective mass scale:**

$$M \gtrsim 1.5 \text{ TeV} \times \sqrt{c_\mu}$$

# Star cooling

Effective coupling of  $\gamma'$  to electrons:  $g_{\gamma'e} = \frac{c_e}{M^2} m_e^2$

Red giant stars:  $\gamma'$  emission via Bremsstrahlung & Compton-like processes

$$g_{\gamma'e}^2 / 4\pi < 2.5 \times 10^{-27} \quad \Rightarrow \quad \frac{M}{\sqrt{c_e}} \gtrsim 3.2 \text{ TeV}$$

For supernovae:  $\nu$  emission rate  $\gg \gamma'$  emission rate  
 $\Rightarrow$  no useful bound on electron- $\gamma'$  coupling  
(strong bound on quark- $\gamma'$  couplings)

# Flavor-changing neutral currents

**Chirality-flip transition:**

$$\Gamma(\mu \rightarrow e\gamma') = c_{e\mu}^2 \frac{m_\mu^5}{8\pi M^4}$$

**Standard model:**

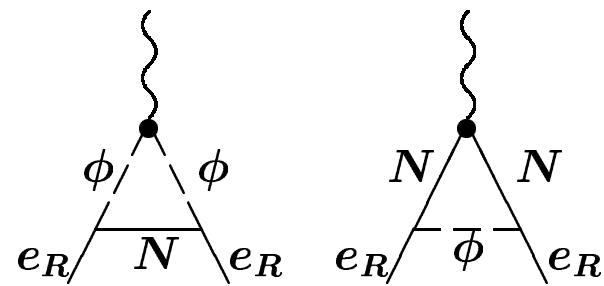
$$\Gamma(\mu \rightarrow e\nu\bar{\nu}) = \frac{m_\mu^5 G_F^2}{192\pi^3} \approx 3.2 \times 10^{-10} \text{ eV}$$

$$\text{Br}(\mu \rightarrow e\gamma') < 3 \times 10^{-5} \quad \Rightarrow \quad \frac{M}{\sqrt{c_{e\mu}}} \gtrsim 15 \text{ TeV}$$

	$l_L$	$e_R$	$N_L$	$N_R$	$H$	$\phi$
$SU(2)_W$	2	1	1	1	2	1
$U(1)_Y$	-1	-2	0	0	+1	-2
$U(1)_D$	0	0	+1	+1	0	-1

**Yukawa interaction:**  $\lambda_\phi^j \bar{e}_R^j N_L \phi + \text{h.c.}$

**Contribution to the  $iP_{\mu\nu} \bar{e}_R \gamma^\mu \partial^\nu e_R$  operator ( $\sim P_{\mu\nu} \bar{l}_L \sigma^{\mu\nu} e_R H$ ):**



$$\Rightarrow (C_e)_{ij} = \frac{g_p}{192\pi^2} \sum_k \left( \lambda_\phi^i \lambda_\phi^{*k} \lambda_e^{kj} + \lambda_e^{ik} \lambda_\phi^k \lambda_\phi^{*j} \right)$$

## Conclusions

New massless gauge boson may couple to quarks and leptons via dimension-6 operators suppressed by the TeV scale!

The lightest particle charged under the new  $U(1)$  is a dark matter candidate.

Look for  $\gamma'$  (coupling to  $t$ -quark) and  $U(1)'$ -charged particles at the Tevatron, LHC, ILC, ...