

# Neutrino Theory at Fermilab

- **Scientists:** Boris Kayser, Stephen Parke, Chris Quigg
- **Post Doc:** Mu-Chen Chen, Olga Mena
- **Users:** Andre de Gouvea and Carl Albright

# Nu Research Papers:

Neutrino Coannihilation on Dark-Matter Relics?

[Gabriela Barenboim \(Valencia U.\)](#), [Olga Mena Requejo](#), [Chris Quigg \(Fermilab\)](#) . FERMILAB-PUB-06-050-T, Apr 2006. 5pp.

e-Print Archive: [astro-ph/0604215](#)

What fraction of boron-8 solar neutrinos arrive at the earth as a  $\nu(2)$  mass eigenstate?

[Hiroshi Nunokawa \(Rio de Janeiro, Pont. U. Catol.\)](#), [Stephen J. Parke \(Fermilab\)](#), [Renata Zukanovich Funchal \(Sao Paulo U.\)](#) . FERMILAB-PUB-05-049-T, Jan 2006. 23pp.

Dedicated to the memory of John Bahcall who championed solar neutrinos for many lonely years.

e-Print Archive: [hep-ph/0601198](#)

Super-NO $\nu$ A: A Long-baseline neutrino experiment with two off-axis detectors.

[Olga Mena Requejo \(Fermilab\)](#), [Sergio Palomares-Ruiz \(Vanderbilt U.\)](#), [Silvia Pascoli \(CERN\)](#) . CERN-PH-TH-2005-050, FERMILAB-PUB-05-050-T, Apr 2005. 17pp.

Published in *Phys.Rev.D72:053002,2005*

e-Print Archive: [hep-ph/0504015](#)

Determining the neutrino mass hierarchy and CP violation in NoVA with a second off-axis detector.

[Olga Mena \(Fermilab\)](#), [Sergio Palomares-Ruiz \(Vanderbilt U.\)](#), [Silvia Pascoli \(CERN & Durham U., IPPP\)](#) . CERN-PH-TH-2005-195, IPPP-05-63, DCPT-05-126, FERMILAB-PUB-05-461-T, Oct 2005. 20pp.

Published in *Phys.Rev.D73:073007,2006*

e-Print Archive: [hep-ph/0510182](#)

Physics potential of the Fermilab NuMI beamline.

[Olga Mena](#), [Stephen J. Parke \(Fermilab\)](#) .

FERMILAB-PUB-05-196-T, May 2005. 22pp.

Published in *Phys.Rev.D72:053003,2005*

e-Print Archive: [hep-ph/0505202](#)

Theory of neutrinos: A White paper.

[R.N. Mohapatra et al.](#) FERMILAB-TM-2342-T, SLAC-PUB-11622, Oct 2005. 143pp.

e-Print Archive: [hep-ph/0510213](#)

Neutrino physics.

[Boris Kayser \(Fermilab\)](#) . SSI-2004-L004, FERMILAB-PUB-05-236-T, Jun 2005. 21pp.

Lectures given at 32nd SLAC Summer Institute on Particle Physics (SSI 2004): Natures Greatest Puzzles, Menlo Park, California, 2-13 Aug 2004.

Published in *eConf C040802:L004,2004*

e-Print Archive: [hep-ph/0506165](#)

## Conference Proceedings:

many ...

## Committees:

Kayser; APS(hep/nucl), NuSAG, Fermilab PAC, P5

## Academic Lectures:

Boris Kayser (4) and Stephen Parke (4)

## Studies:

Proton Driver: Parke and Mena

## FNAL Experiments:

MINOS, mini-BOONE

NOvA

# Hierarchy Determination: Counting Expts at First Osc. Max.

- Neutrino  $\nu$  Anti-Neutrino **NO $\nu$ A** Expt.
- Neutrino  $\nu$  Neutrino Two Expts Different L's and EQUAL E/L's: **NO $\nu$ A+T2K**
- Neutrino  $\nu$  Anti-Neutrino Two Expts Different L's: **NO $\nu$ A+T2K**

T2K

JHF  $\rightarrow$  Super-Kamiokande

- 295 km baseline
- Super-Kamiokande:
  - 22.5 kton fiducial
  - Excellent  $e/\mu$  ID
  - Additional  $\pi^0/e$  ID
- Hyper-Kamiokande
  - 20 $\times$  fiducial mass of SuperK
- Matter effects small
- Study using fully simulated and reconstructed data

The NUMI Beamline

NO $\nu$ A

Two functionally identical neutrino detectors

Fermilab 10 km Soudan

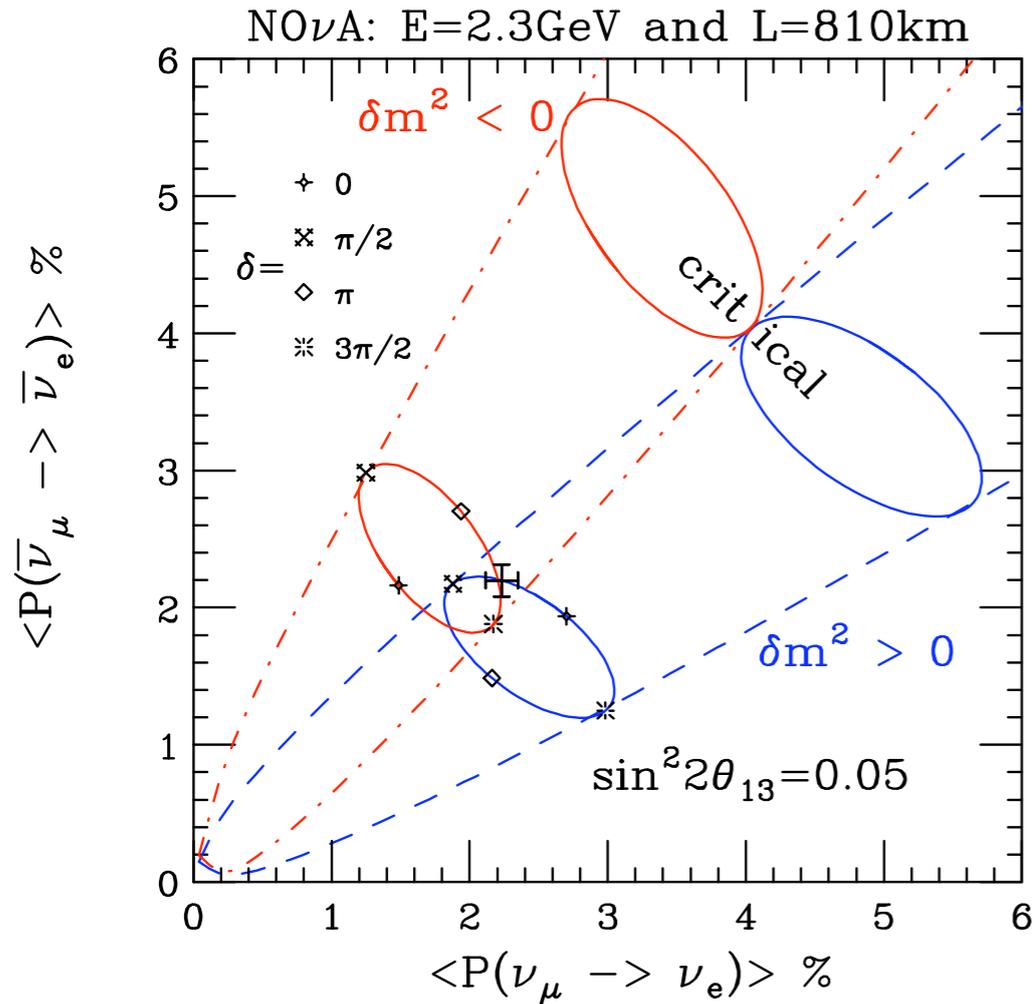
730 km 12 km

Det. 1 Det. 2

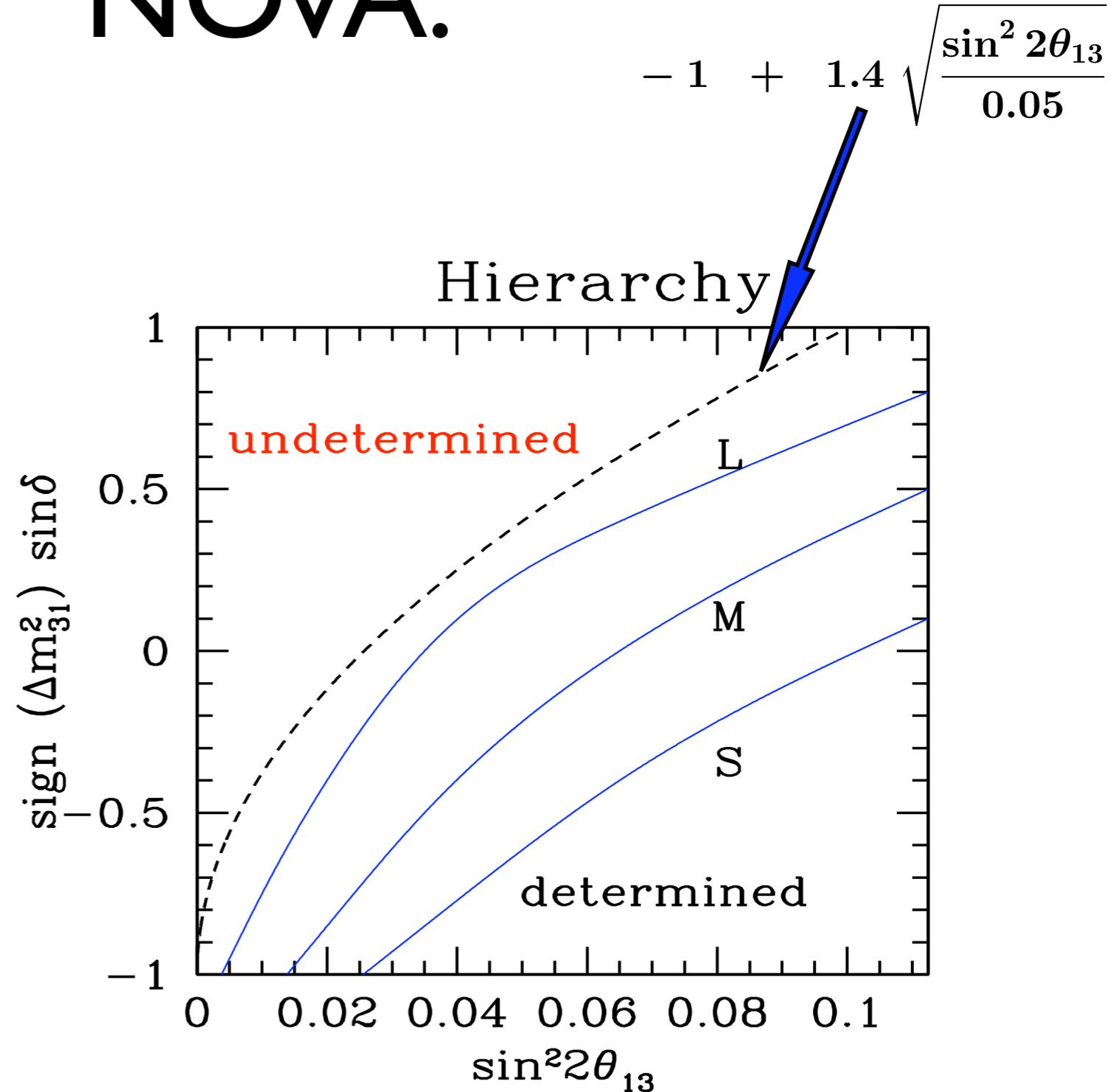
# I

## PHYSICS POTENTIAL OF THE FERMILAB NUMI BEAMLINE

Olga Mena, Stephen J. Parke [hep-ph/0505202](#)



## NO $\nu$ A:



S: 4 +4 yrs

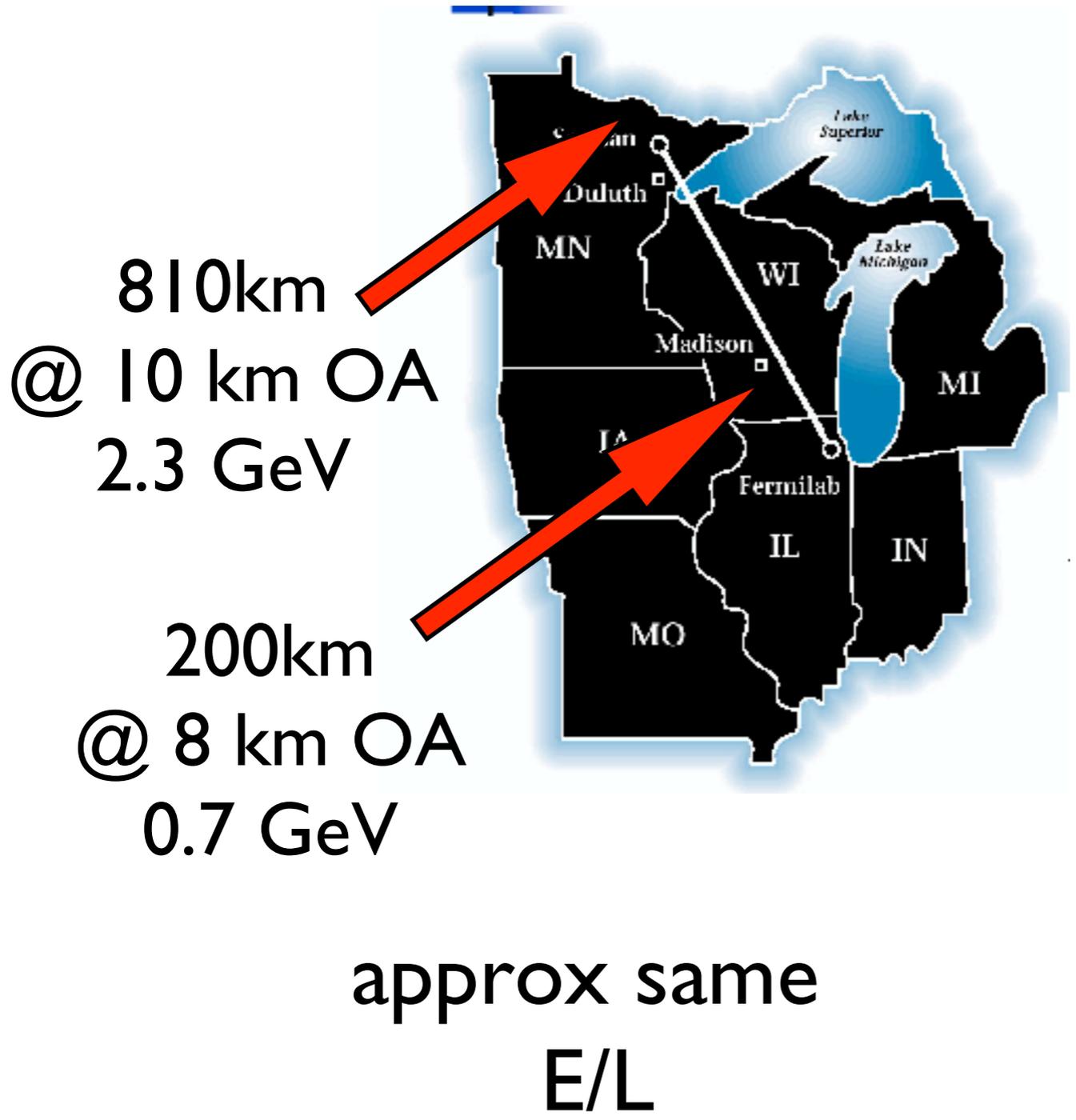
M (=5\*S):

Proton Driver

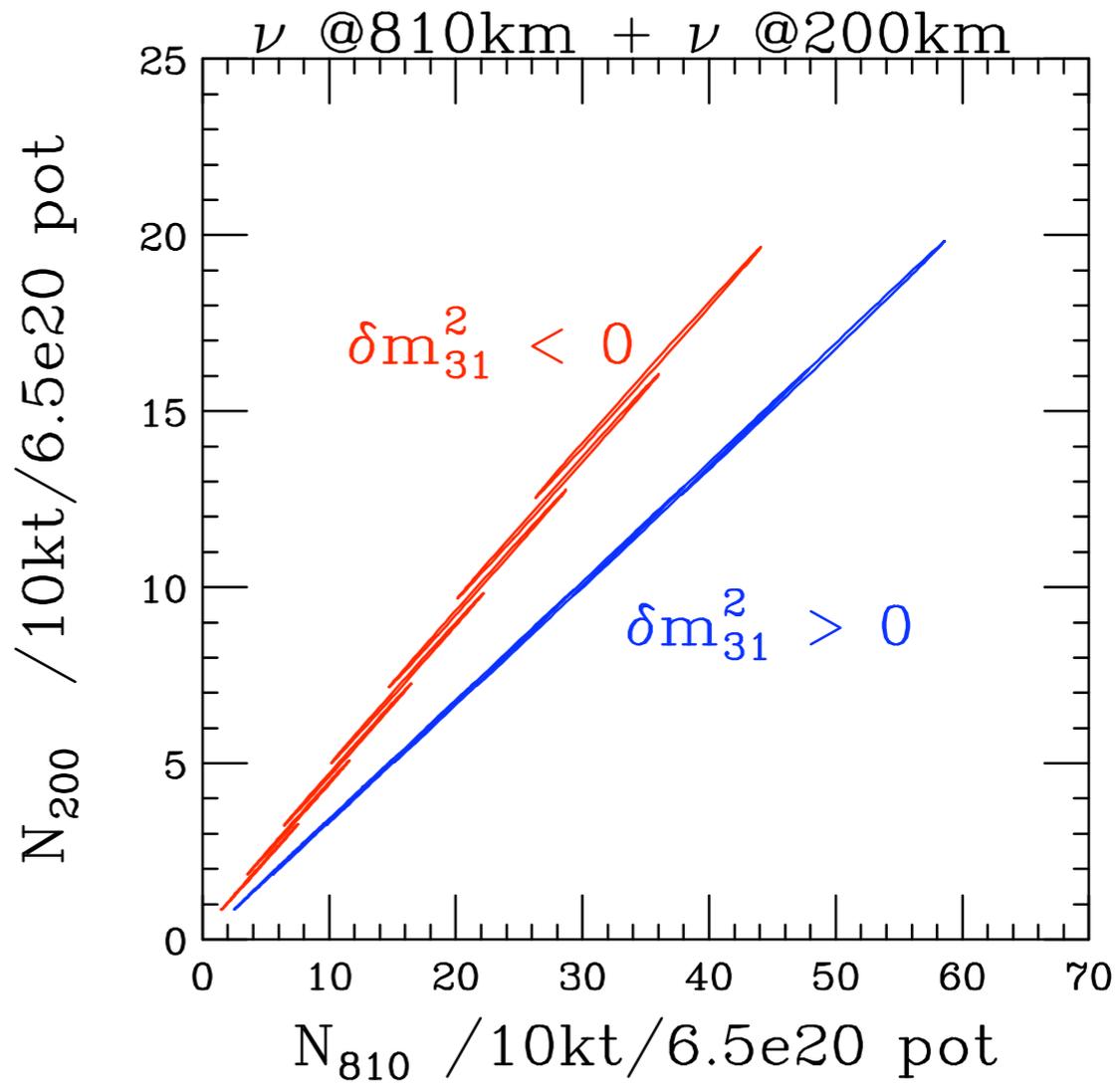
L (=5\*M):

PD + Liquid Argon

**NOvA plus “NEAR” DETECTOR**

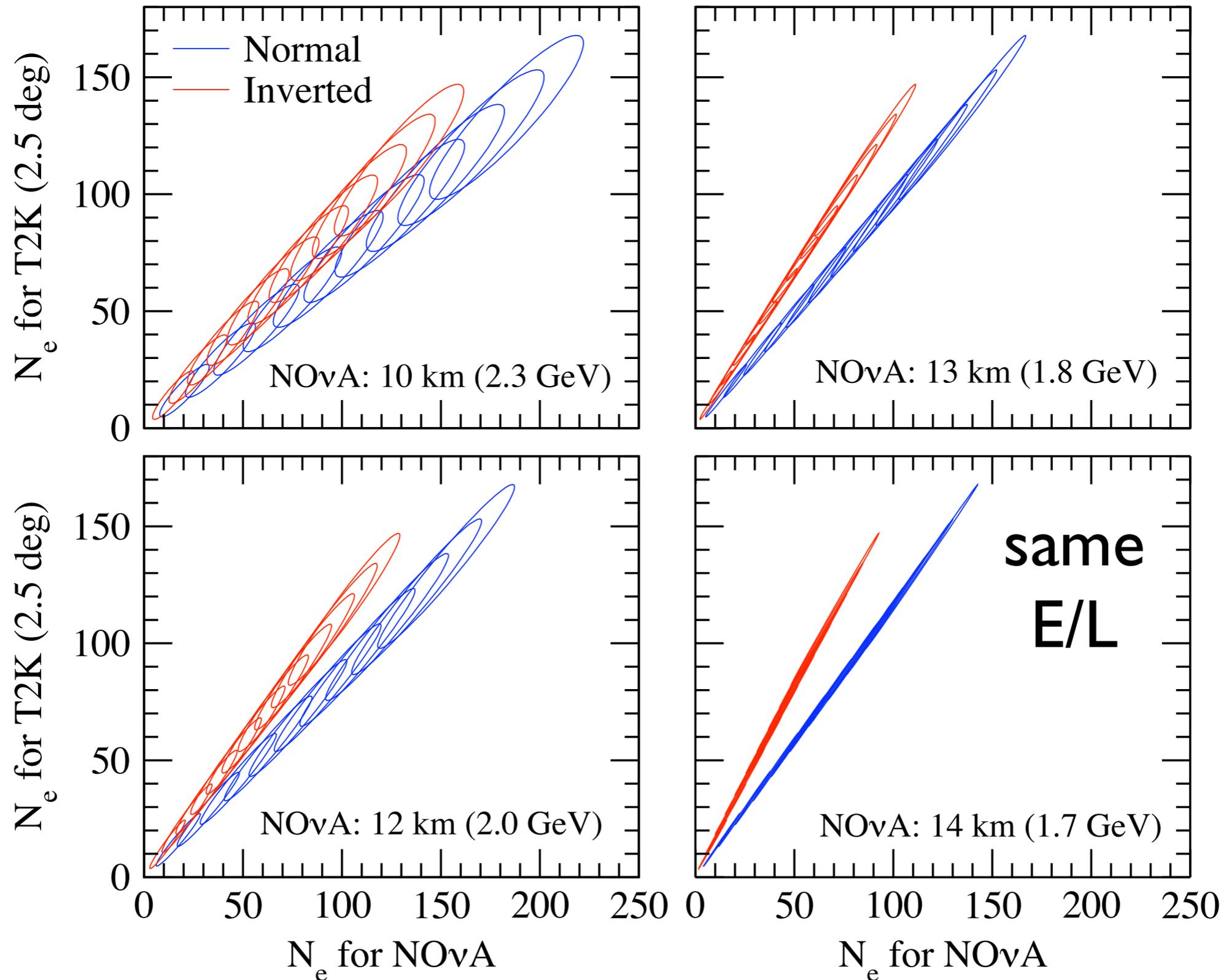


**Neutrino - Neutrino**



$\sin^2 2\theta_{13} = (1, 2, 3, 4.3, 6, 7.4, 9.5) \cdot 10^{-2}$

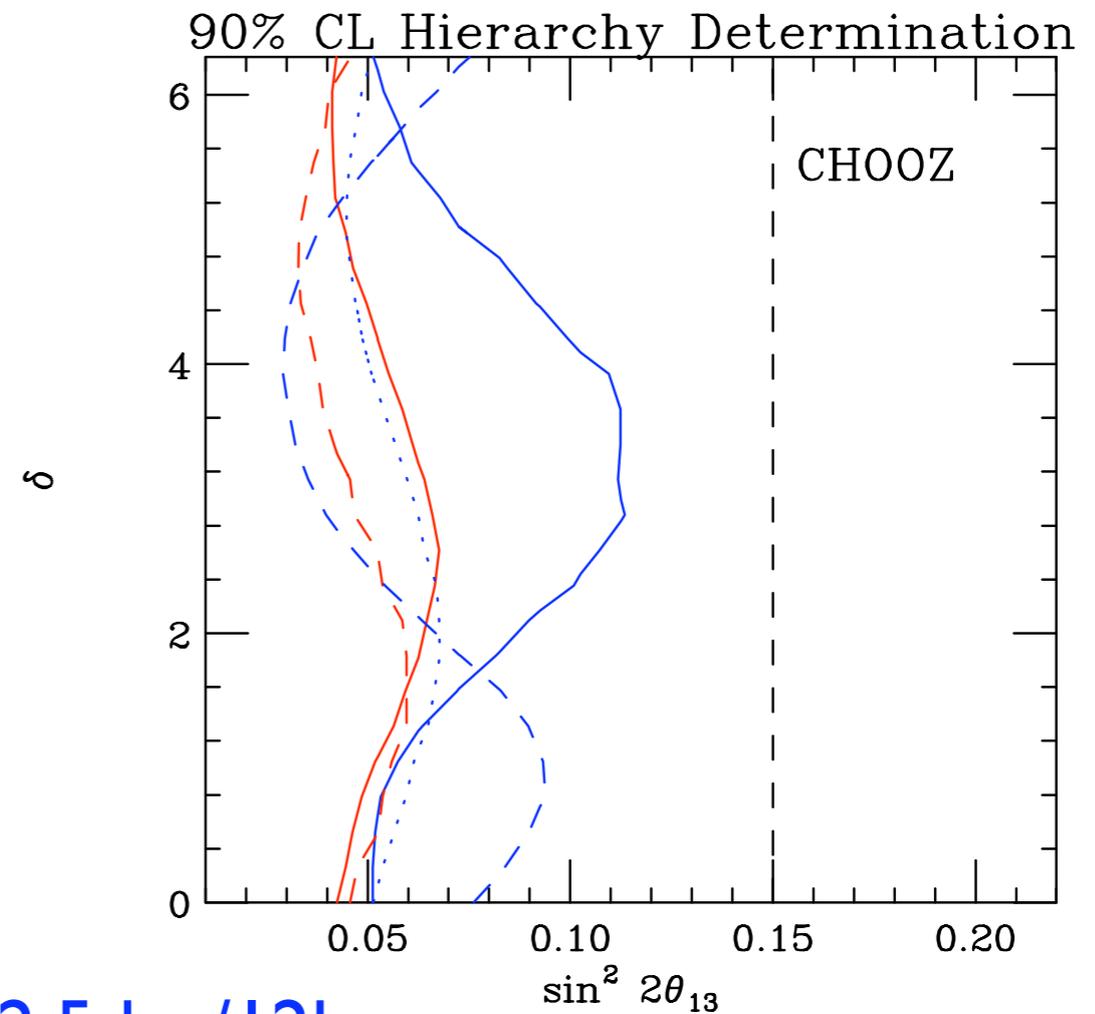
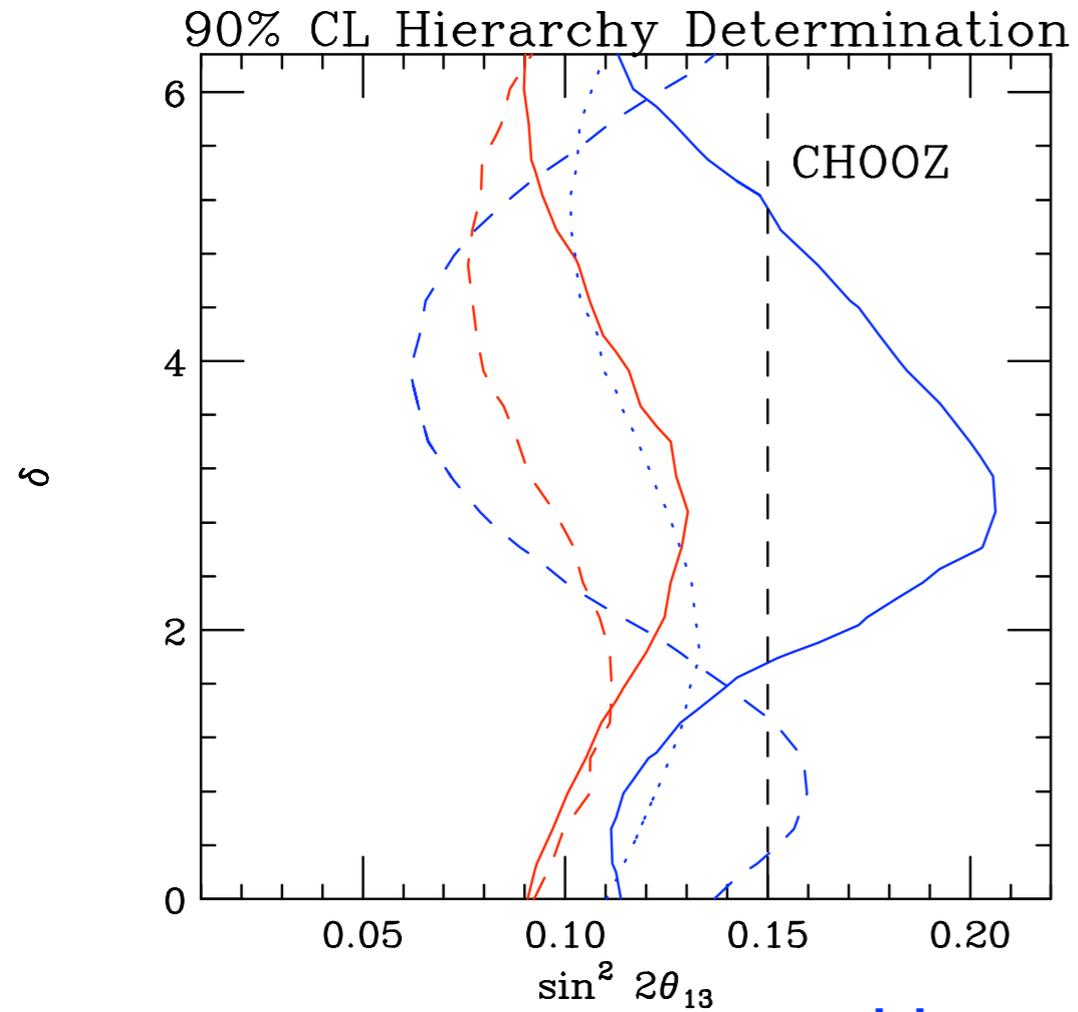
# What about combining T2K and NOvA? Neutrinos Only



# T2K (deg) + NOvA (km) Off Axis:

## phase I

## phase II



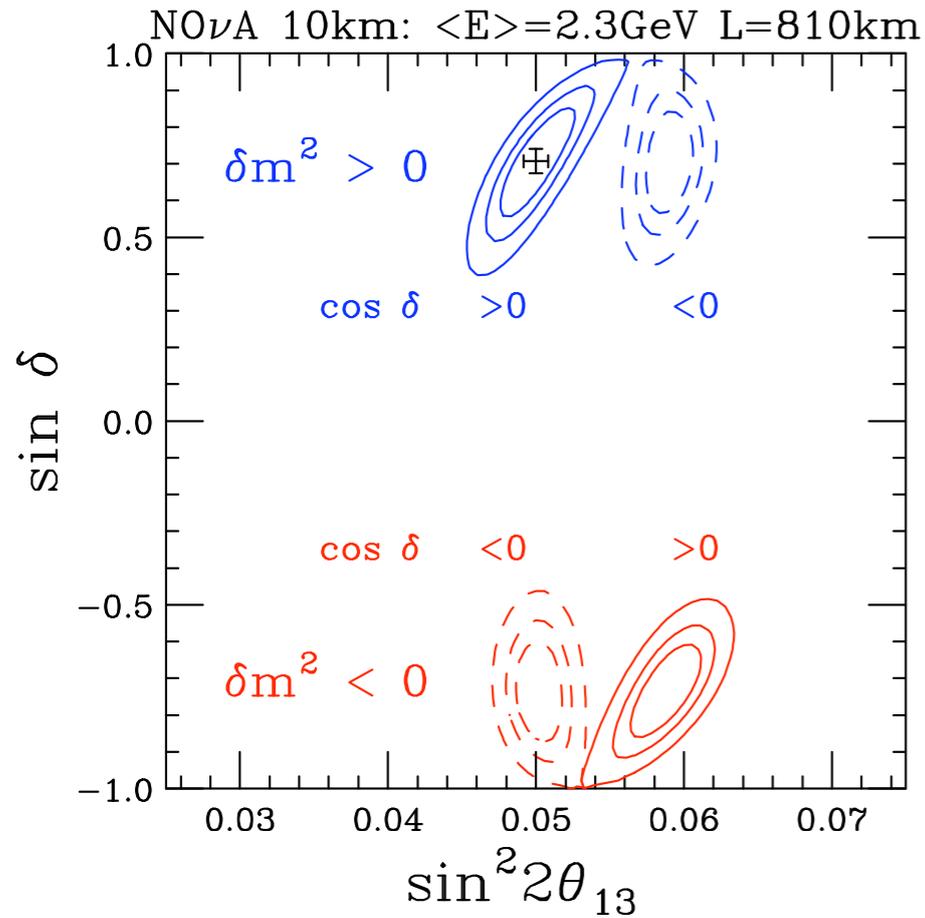
blue solid 2.5deg/12km  
red solid 2.5deg/13km  
red dashed 2.5deg/14km  
blue dashed 2.5 deg/16km  
blue dotted 2.0deg/12km

# III

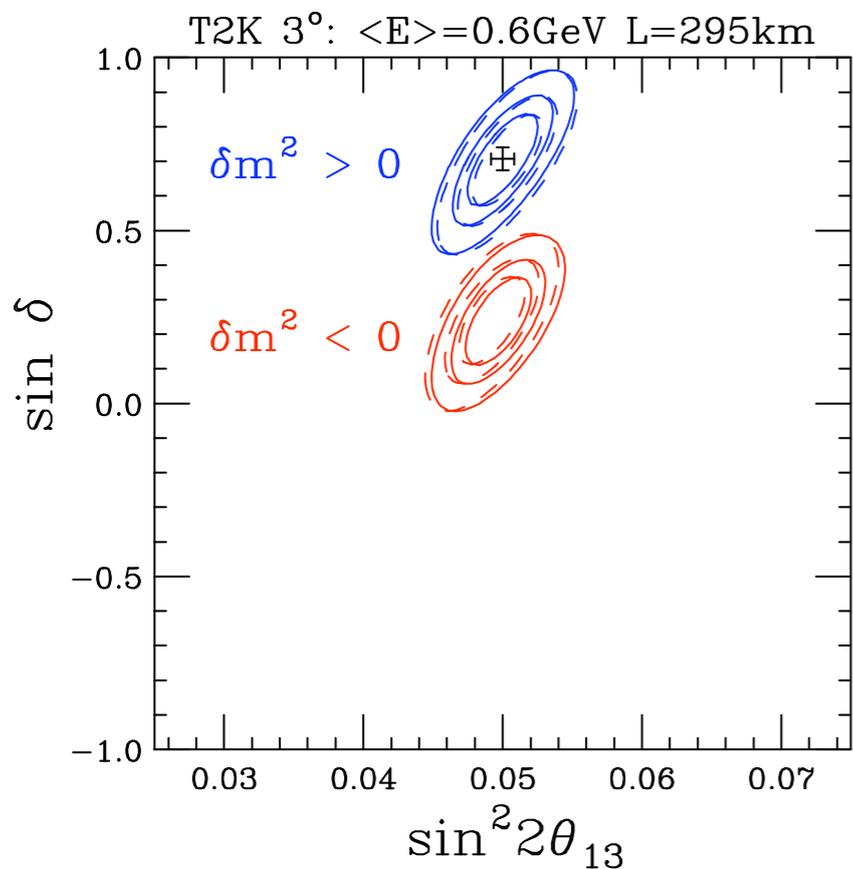
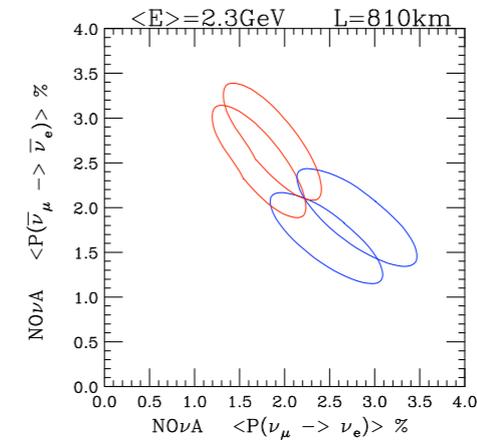
## UNTANGLING CP VIOLATION AND THE MASS HIERARCHY IN LONG BASELINE EXPERIMENTS.

Olga Mena, Stephen J. Parke

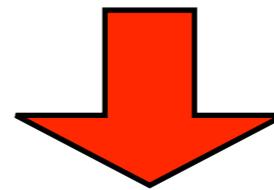
hep-ph/0408070



$$\langle \sin \delta \rangle_+ - \langle \sin \delta \rangle_- = 1.41 \sqrt{\frac{\sin^2 2\theta_{13}}{0.05}} \quad \text{for NO}\nu\text{A.}$$



$$\langle \sin \delta \rangle_+ - \langle \sin \delta \rangle_- = 0.47 \sqrt{\frac{\sin^2 2\theta_{13}}{0.05}} \quad \text{for T2K}$$



$$| \langle \sin \delta \rangle_{fake}^{T2K} - \langle \sin \delta \rangle_{fake}^{NO\nu A} | = 0.94 \sqrt{\frac{\sin^2 2\theta_{13}}{0.05}}$$

For an “early” determination of the  
Neutrino Mass Hierarchy,

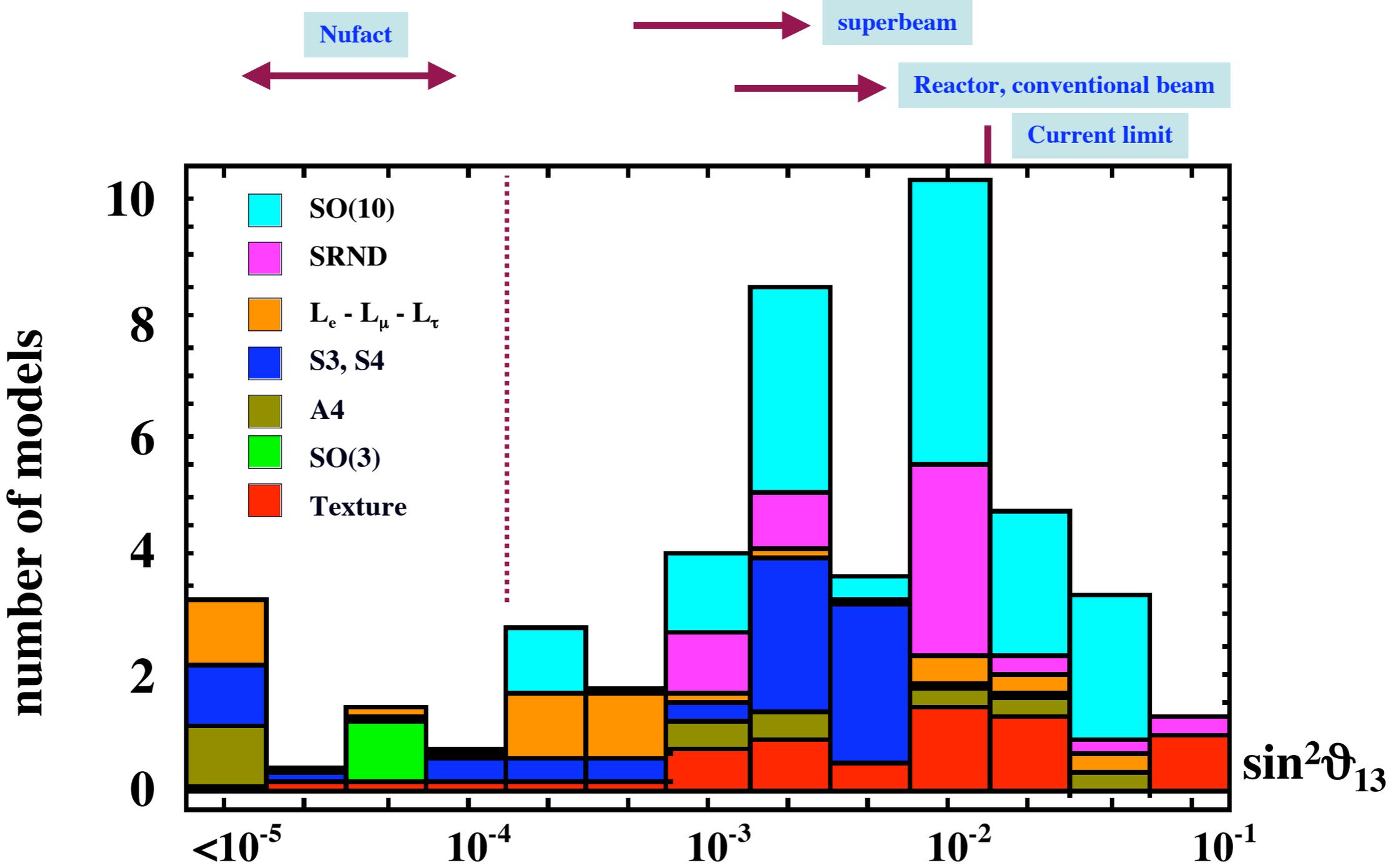
NO $\nu$ A

is an Essential Ingredient.

# Albright & Chen:

Predictions of all 46 models for 1-3 mixing:

C. Albright & M.-C.C. (2006)



# What Fraction of $^8\text{B}$ Solar Neutrinos arrive at the Earth as a $\nu_2$ mass eigenstate?<sup>1</sup>

Stephen Parke, Hiroshi Nunokawa & Renata Zukanovich-Funchal  
hep-ph/0601198

$$\delta m_{\odot}^2 = 8.0 \times 10^{-5} eV^2$$

$$\sin^2 \theta_{\odot} = 0.31$$

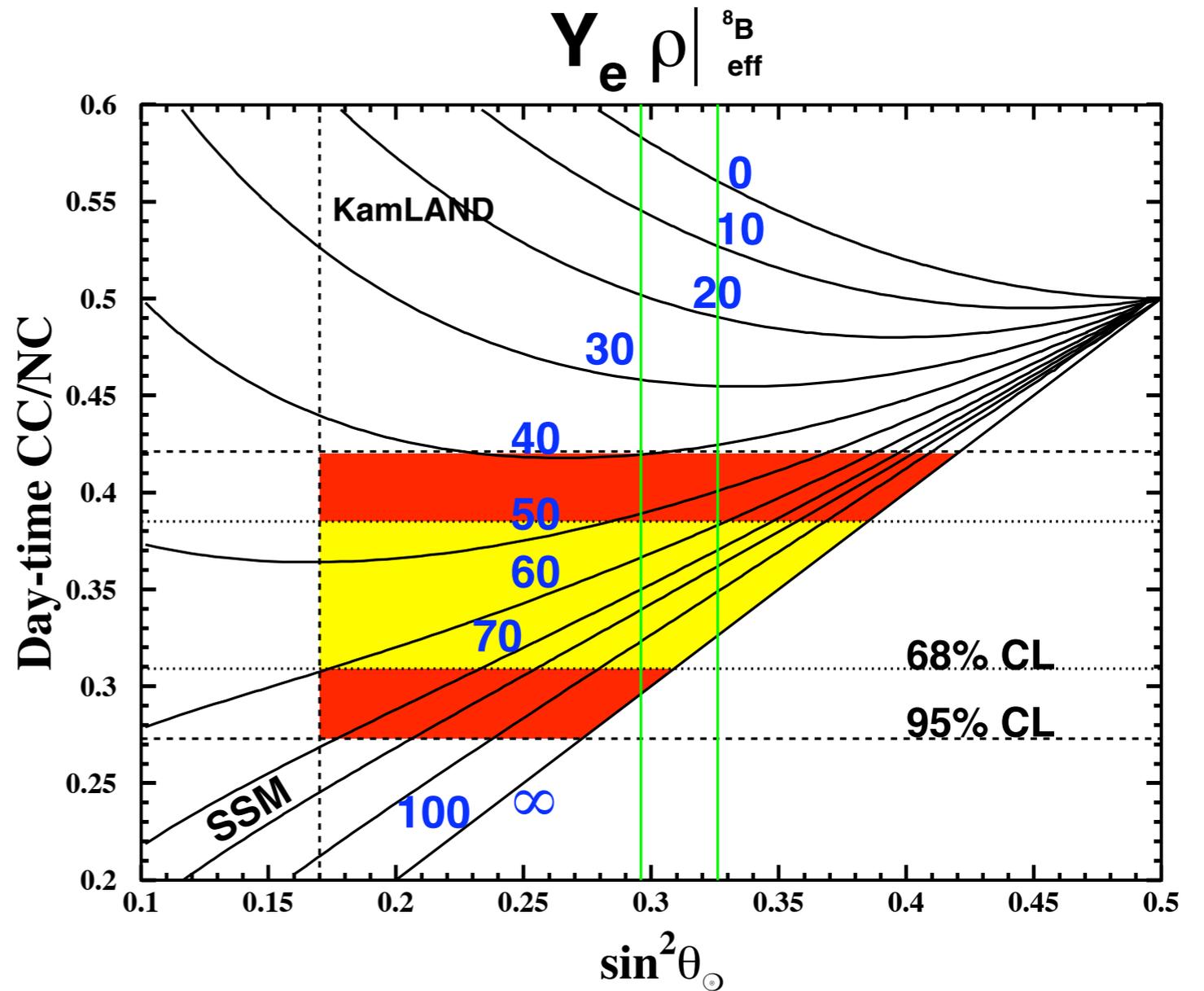
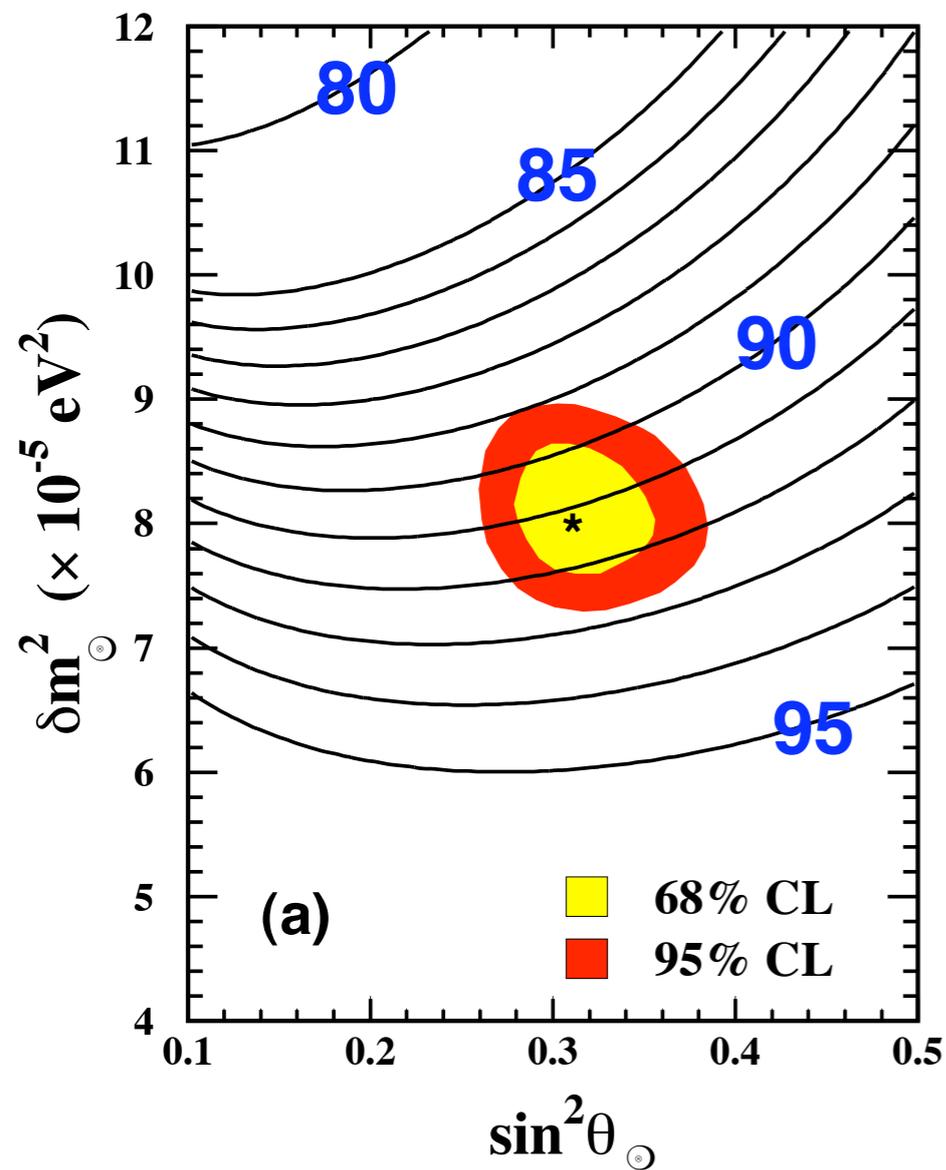
$$\nu_2 \sim \frac{1}{\sqrt{3}} (\nu_e + \nu_{\mu} + \nu_{\tau})$$

$$L_{osc} = \frac{4\pi E}{\delta m_{\odot}^2} = 300 km \left( \frac{E}{10 MeV} \right) \quad \text{OR} \quad \Delta \equiv \frac{\delta m_{\odot}^2 L}{4E} = 10^6 \left( \frac{10 MeV}{E} \right)$$

**Effectively Incoherent !!!**

# What Fraction of $^8\text{B}$ Solar Neutrinos arrive at the Earth as a $\nu_2$ mass eigenstate?<sup>1</sup>

$^8\text{B}$ :  $\nu_2$  fraction (%)



91 +/- 2 %

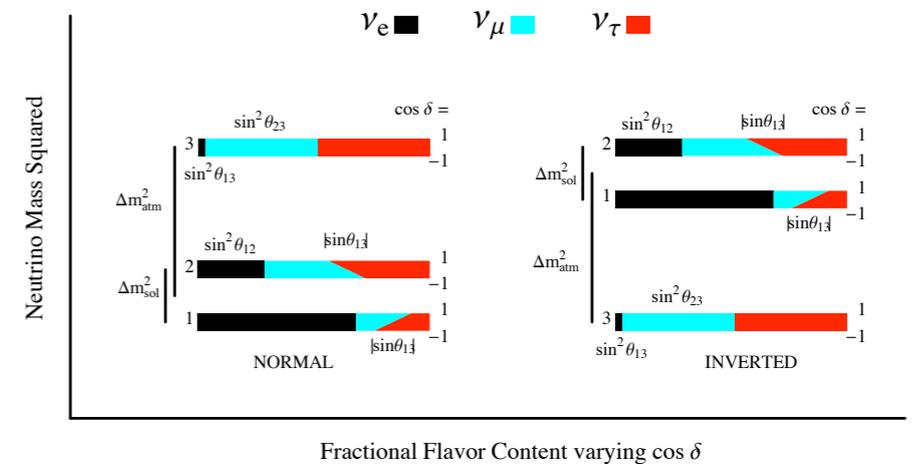
# 3 flavor fractions:

$$\mathcal{F}_1 \approx f_1 = 0.09 \mp 0.02,$$

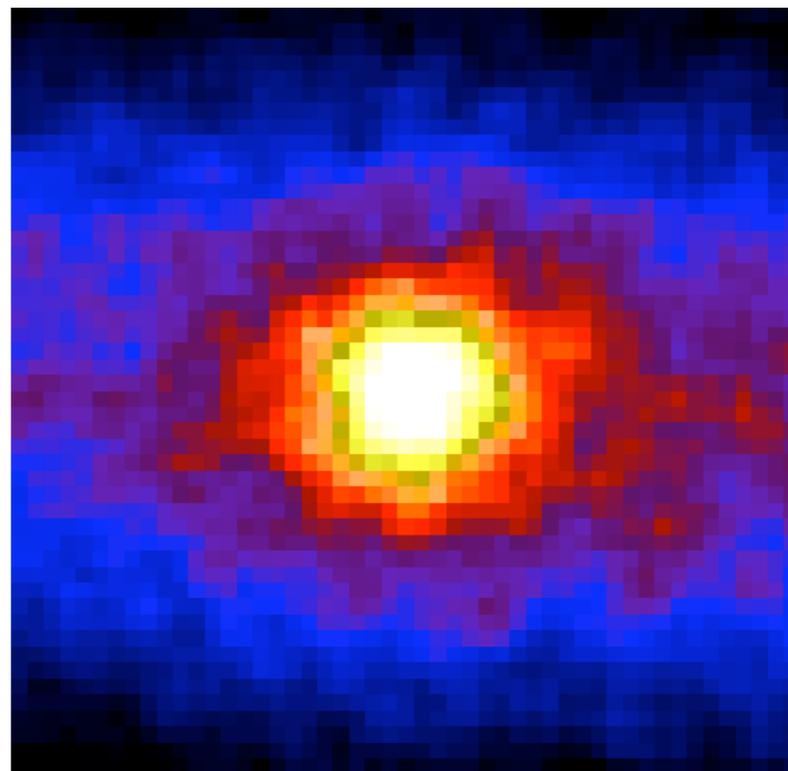
$$\mathcal{F}_2 = f_2 - \sin^2 \theta_{13} \approx 0.91 \pm 0.02 - \sin^2 \theta_{13},$$

$$\mathcal{F}_3 = \sin^2 \theta_{13}.$$

$$|U_{e2}|^2 \approx \sin^2 \theta_{\odot}^8 B + (0.53_{-0.04}^{+0.06}) \sin^2 \theta_{13}.$$



SK



These are  $\nu_2$  Neutrinos !!!

**extras:**

# Neutrino Coannihilation on Dark-Matter Relics?

Gabriela Barenboim,<sup>1,\*</sup> Olga Mena Requejo,<sup>2,†</sup> and Chris Quigg<sup>2,‡</sup>

[astro-ph/0604215](#)

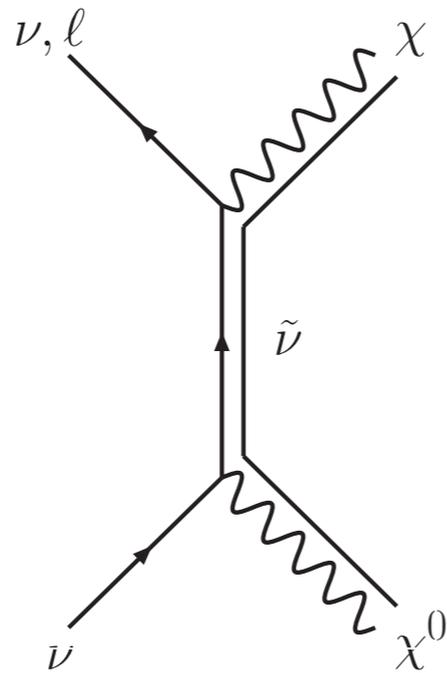


FIG. 1: Resonant sneutrino formation in neutrino–neutralino collisions. Double lines denote superpartners.

detector in the vicinity of Earth

$$1.3 \times 10^{-21} \text{ events cm}^{-2} \text{ y}^{-1}.$$