

# Axion Production and Detection with Superconducting RF Cavities

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**Ryan Janish**

(UC Berkeley)

[RJ, Narayan, Rajendran, Riggins, 1904.07245]

# Axion-like Particles (ALPs)

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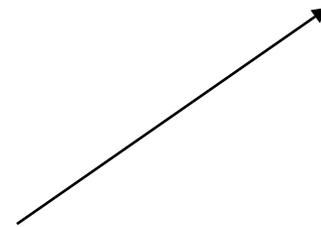
**A new light boson?** Physics in the far UV can lead to light, weakly-coupled particles.

ALPs that couple to photons are a generic possibility

$$\mathcal{L} \supset \frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} (\partial_\mu a)^2 - \frac{1}{2} m_a^2 a^2 + \frac{1}{4} g a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

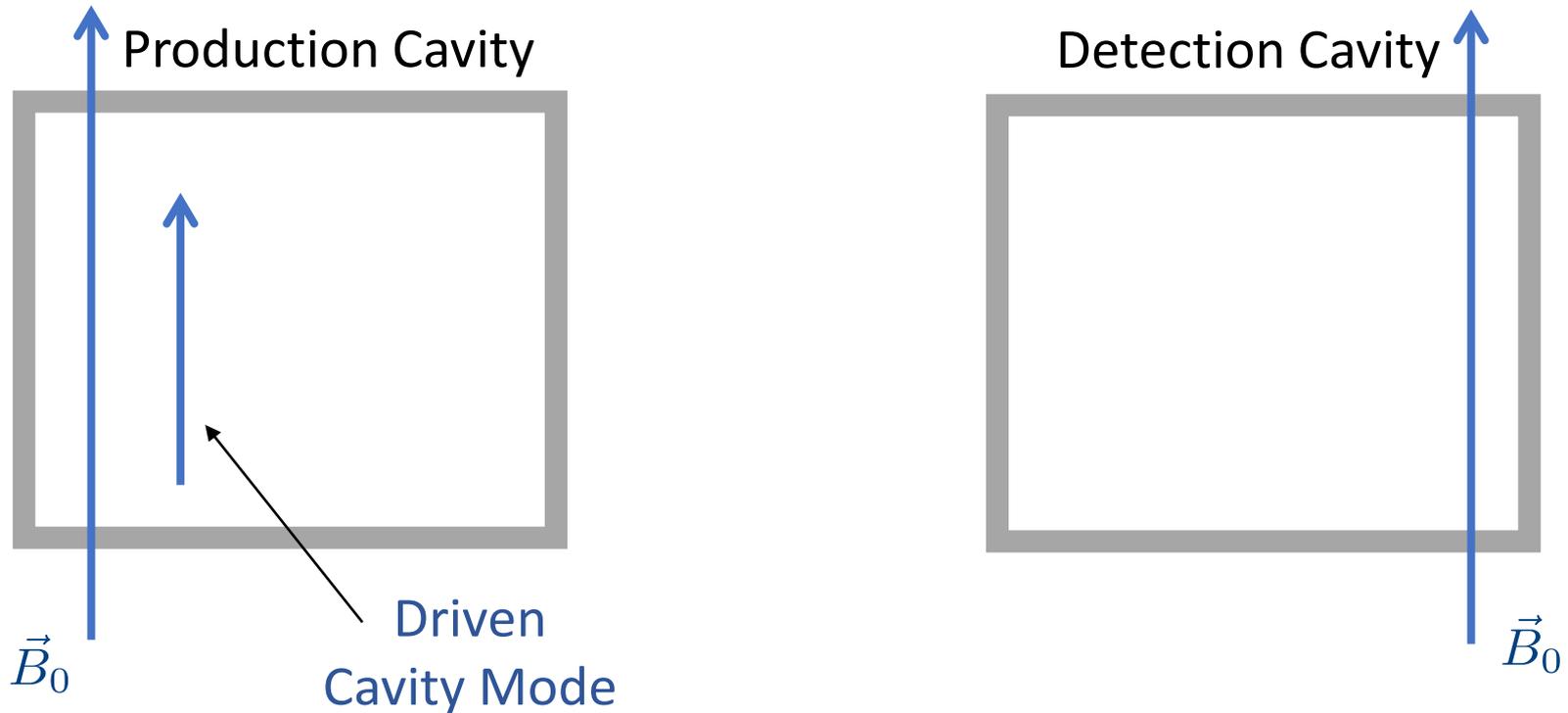
**Detection Opportunity**

ALP-photon mixing in a magnetic field



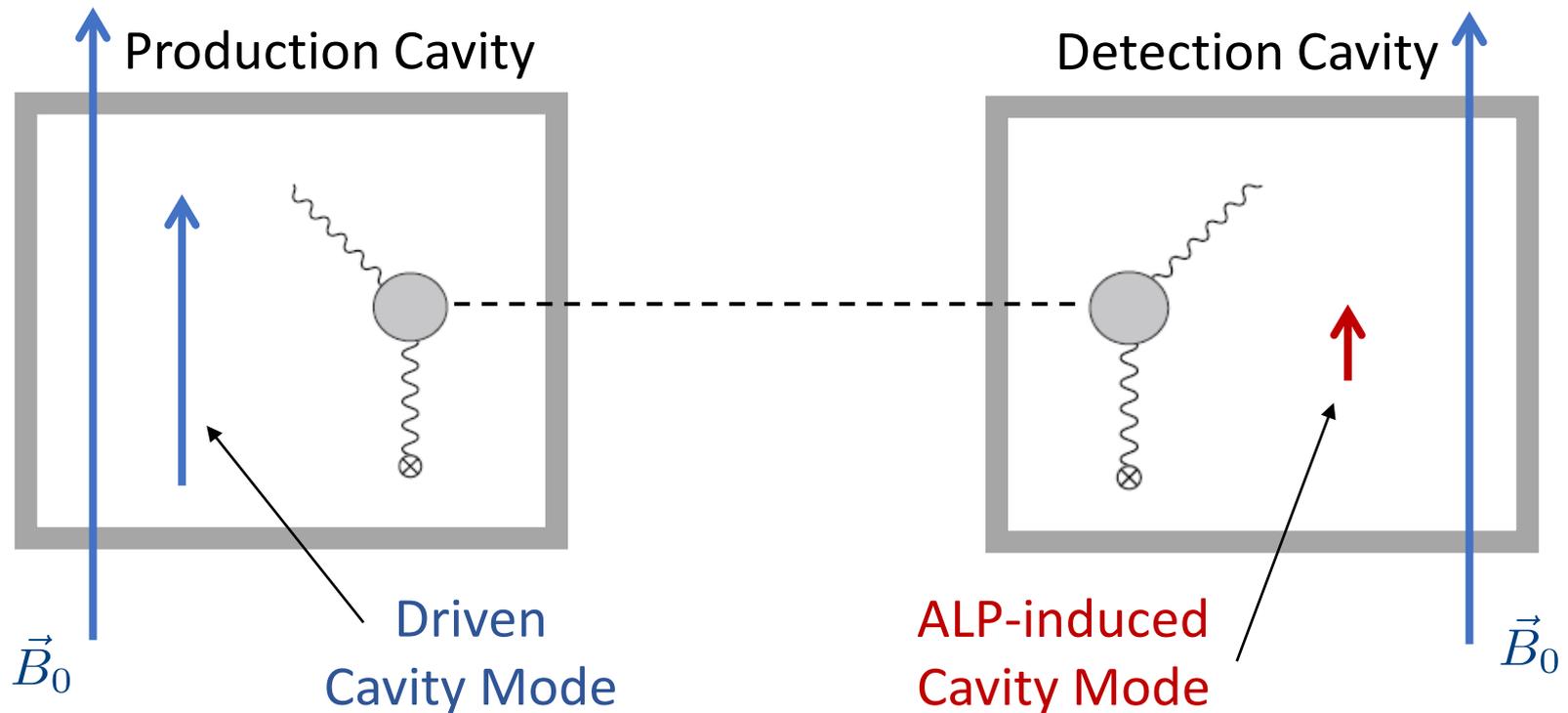
# Light Shining Through Walls (LSW)

[Hoogeveen, '92]



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# Light Shining Through Walls (LSW)

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[Graham et al '16]

## Optical Cavities

$\omega, L^{-1}$  are independent

**ALPS**  $g < 5 \times 10^{-8} \text{ GeV}^{-1}$   
for  $m_a < \text{meV}$

Next generation with  $L \sim 100 \text{ m}$   
ALPS II (projected):  $g < 2 \times 10^{-11} \text{ GeV}^{-1}$

## RF Cavities

$\omega, L^{-1} \sim \mathcal{O}(\text{GHz})$

**CROWS**  $g < 10^{-7} \text{ GeV}^{-1}$   
for  $m_a < \mu\text{eV}$

Next generation ???

# Light Shining Through Walls (LSW)

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Next generation ???

**This work: Utilize superconducting RF technology  
to reach  $g < 7 \times 10^{-12} \text{ GeV}^{-1}$  in a next  
generation LSW ALP search**

See [Bogorad, Hook, Kahn, Soreq, '19] for a different approach

# Light Shining Through Walls (LSW)

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$$P_{\text{signal}} = P_{\text{input}} \left( \frac{gB_0}{\omega} \right)^4 Q_{\text{pc}} Q_{\text{dc}} |G|^2$$

# Light Shining Through Walls (LSW)

---

$$P_{\text{signal}} = P_{\text{input}} \left( \frac{gB_0}{\omega} \right)^4 Q_{\text{pc}} Q_{\text{dc}} \underline{|G|^2}$$

 O(0.1) geometric form factor, exponentially suppressed for  $m_a > \omega$

# Light Shining Through Walls (LSW)

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$$P_{\text{signal}} = P_{\text{input}} \left( \frac{gB_0}{\omega} \right)^4 \frac{Q_{\text{pc}} Q_{\text{dc}} |G|^2}{\phantom{Q_{\text{pc}} Q_{\text{dc}} |G|^2}}$$



Conventional Conducting RF:  $Q \sim 10^5 - 10^6$

Superconducting RF:  $Q \sim 10^{10} - 10^{12}$

# Light Shining Through Walls (LSW)

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$$P_{\text{signal}} = P_{\text{input}} \left( \frac{gB_0}{\omega} \right)^4 \frac{Q_{\text{pc}} Q_{\text{dc}} |G|^2}{Q}$$



Conventional Conducting RF:  $Q \sim 10^5 - 10^6$

Superconducting RF:  $Q \sim 10^{10} - 10^{12}$

## Type-II Superconductor Critical Field O(0.2 T)

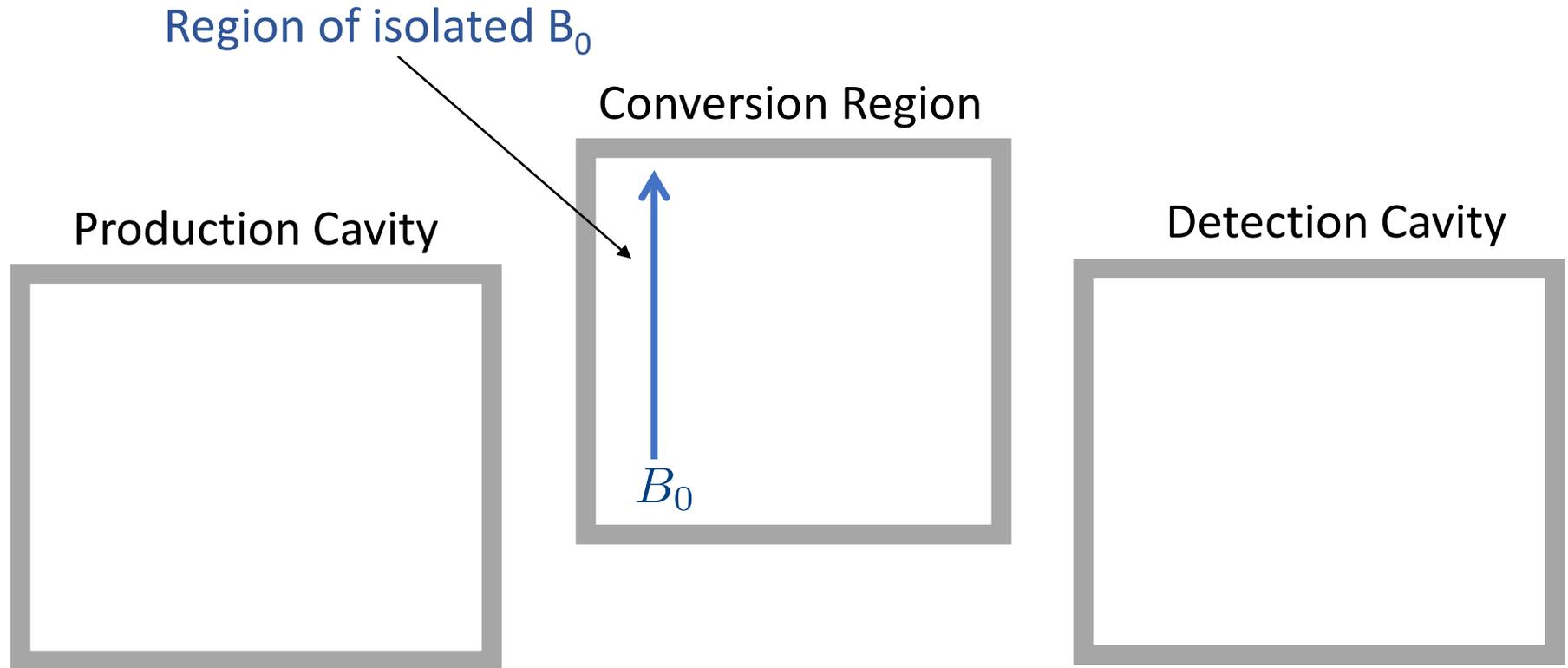
Fields exceeding critical will degrade Q.

Flux tubes penetrate the SC, have a dissipative interaction with an electric current.

Challenge: re-design such that large B and SRF cavity can co-exist

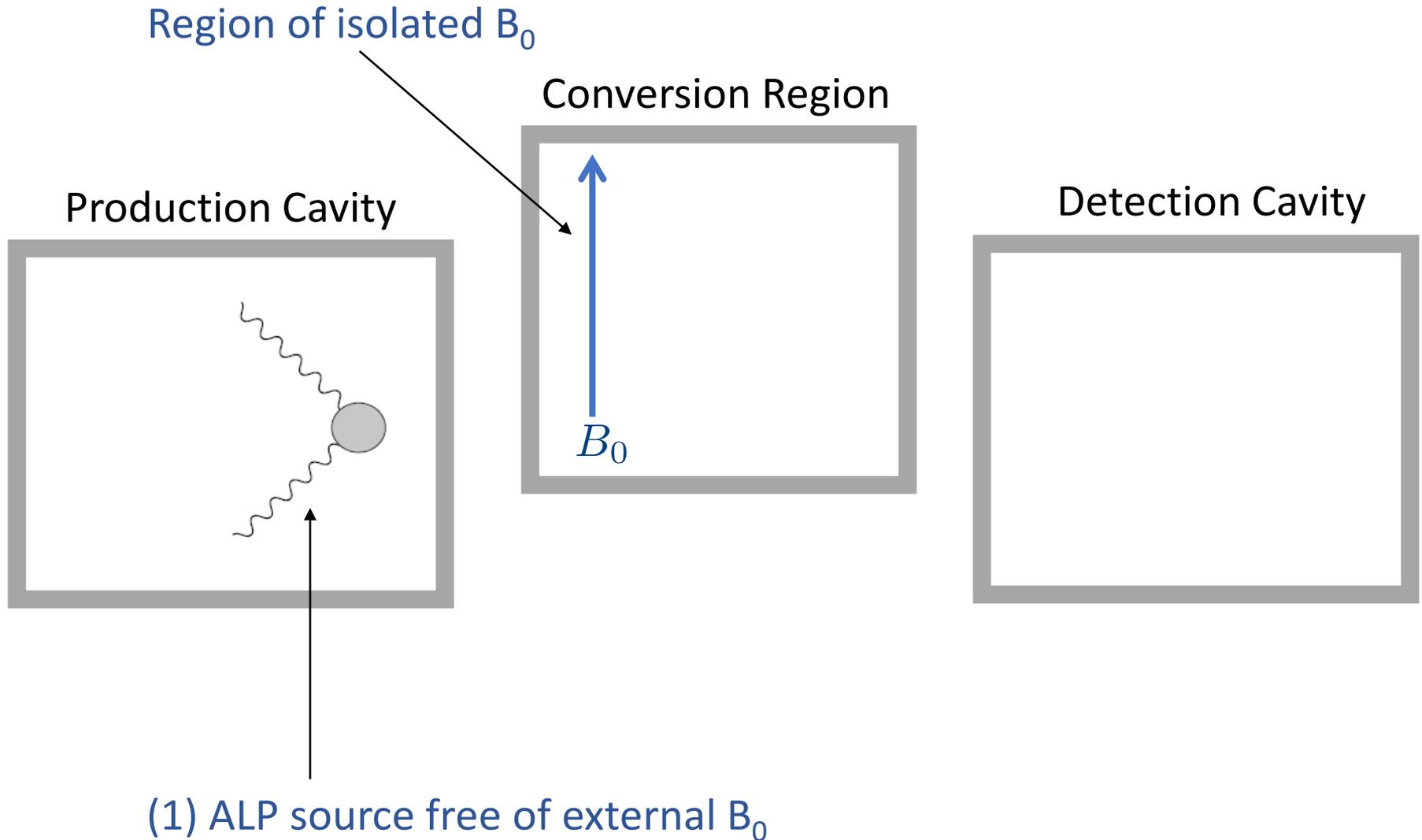
# LSW with SRF Cavities

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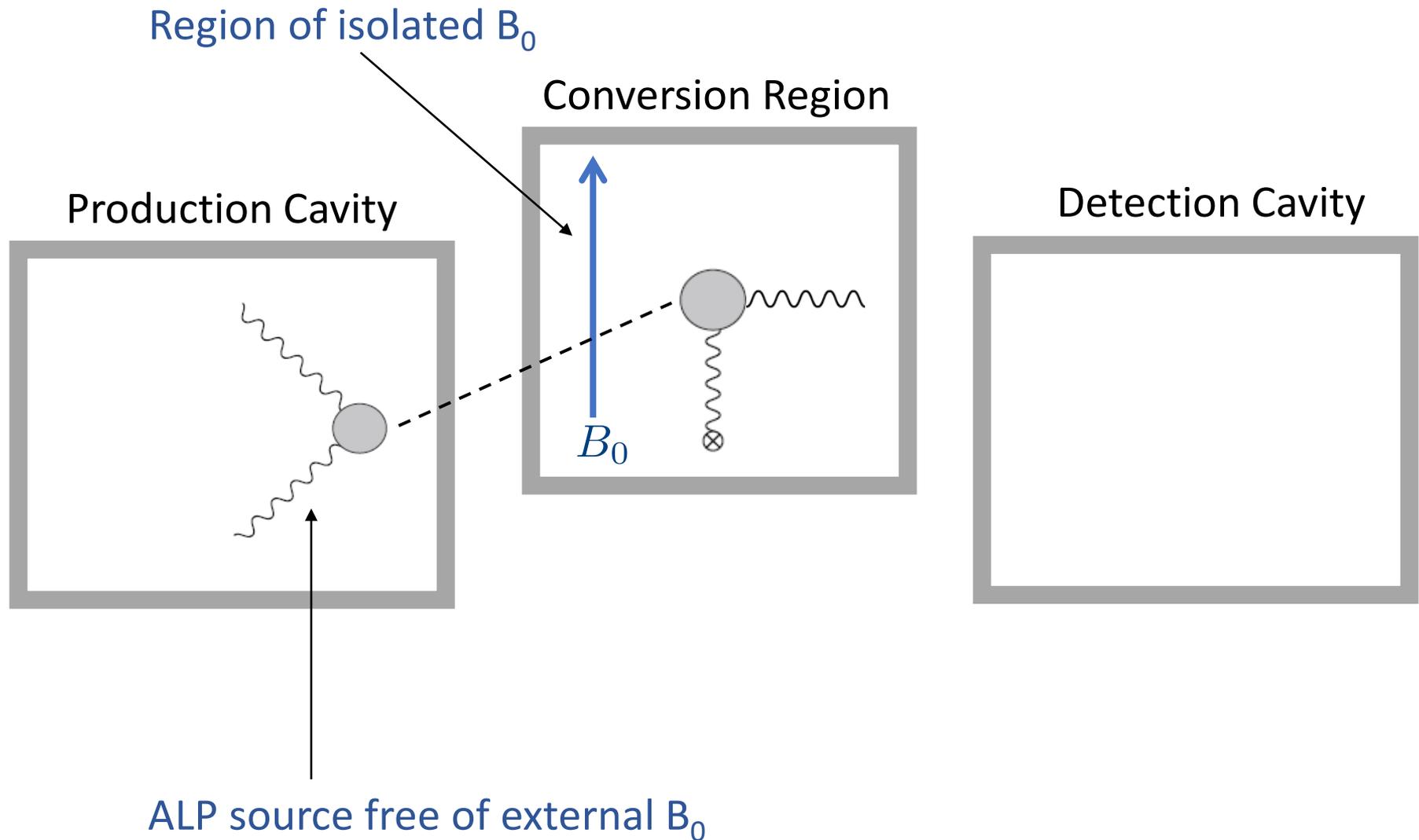
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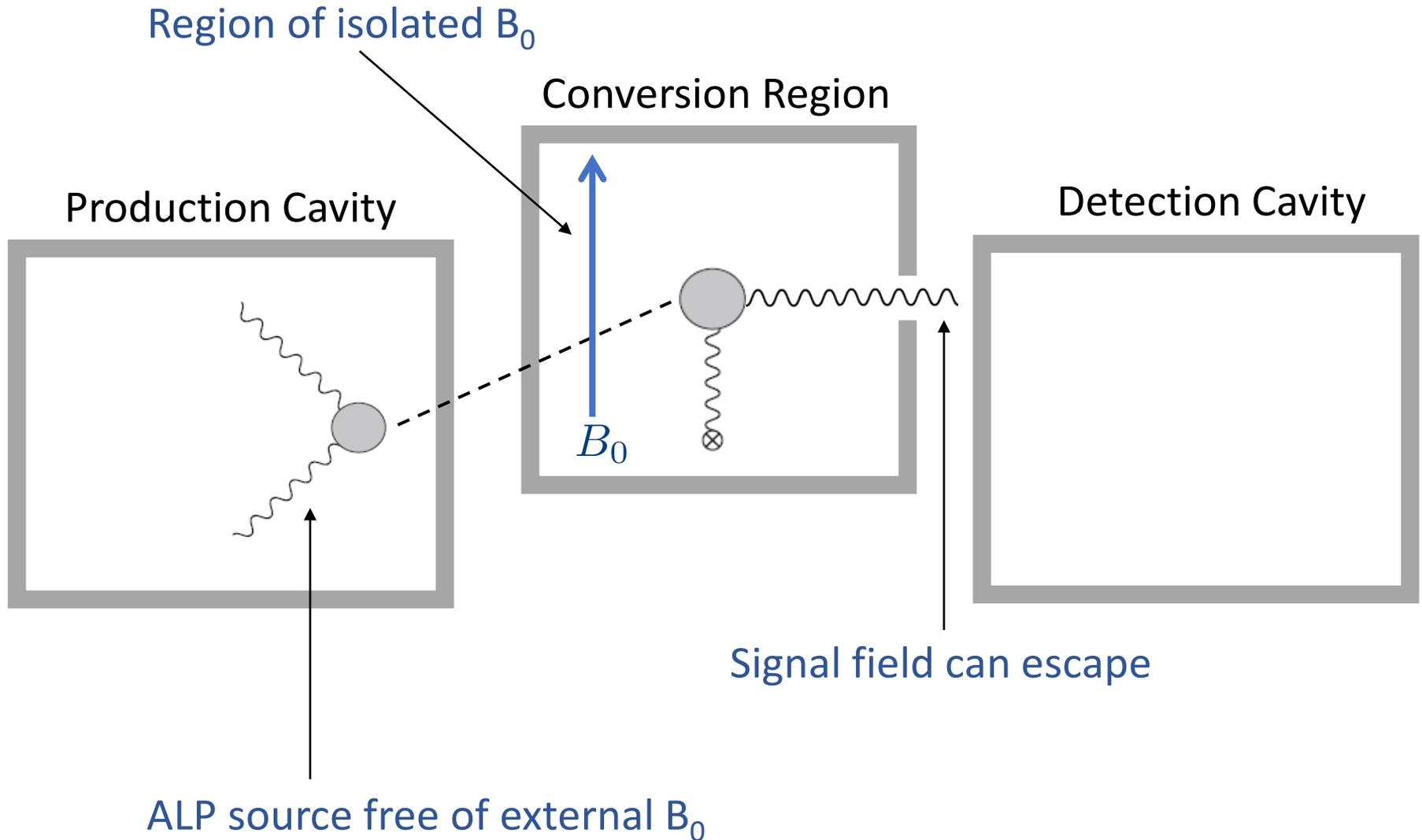


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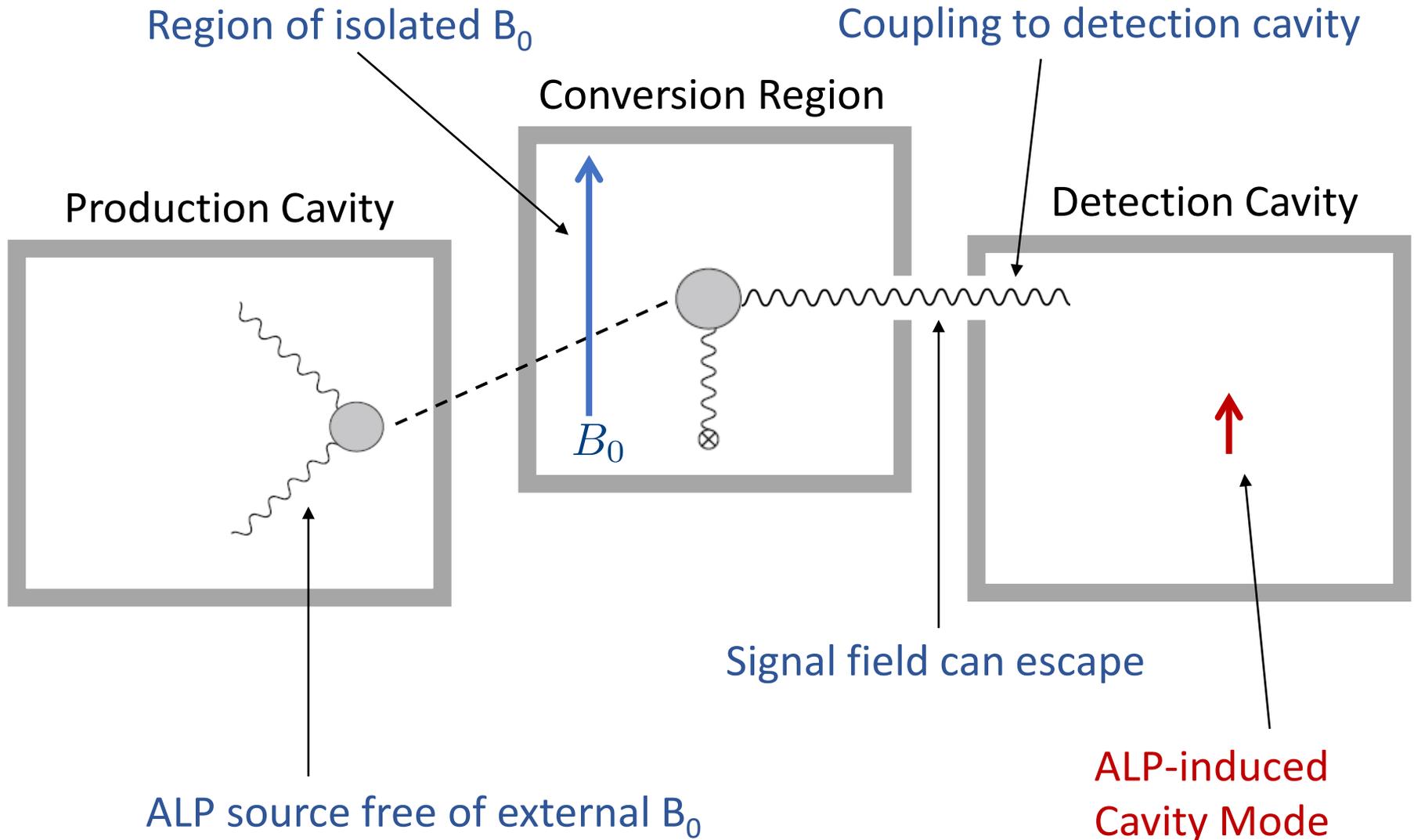
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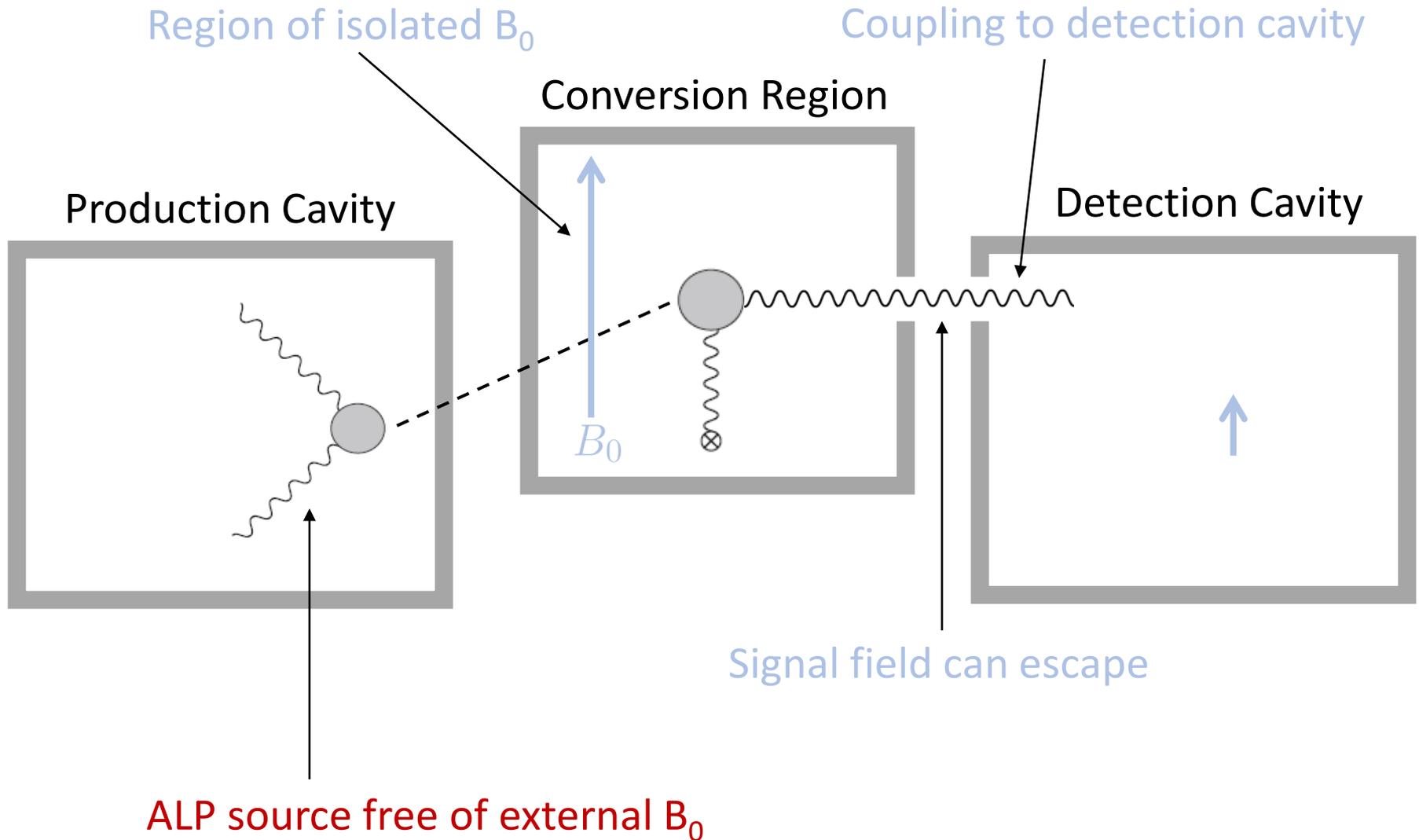
# LSW with SRF Cavities



# LSW with SRF Cavities



# LSW with SRF Cavities



# SRF Axion Source

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## Axion production

$$\text{Equation of Motion: } (\square + m_a^2)a(x) = -g\vec{E} \cdot \vec{B}$$

$$\Rightarrow a(x) = -ge^{i\omega t} \int_{\text{pc}} d^3y \frac{e^{ik|\vec{x}-\vec{y}|}}{4\pi|\vec{x}-\vec{y}|} (\vec{E} \cdot \vec{B})$$

# SRF Axion Source

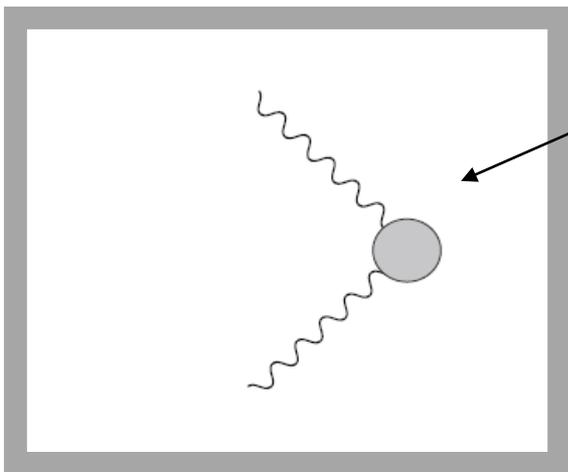
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## ALP source free of external $B_0$



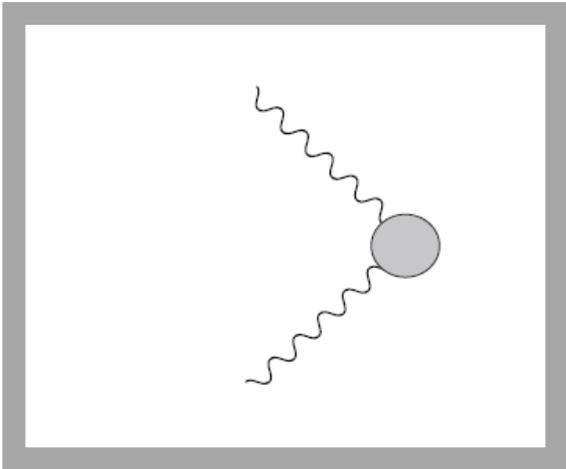
Drive cavity mode(s) such that  $\vec{E} \cdot \vec{B}$  is not identically zero.

May require exciting two modes,  $\omega_1$  and  $\omega_2$ , which produces axions of frequency  $\omega_1 \pm \omega_2$

# SRF Axion Source

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ALP source free of external  $B_0$



Input power is limited to prevent quenching

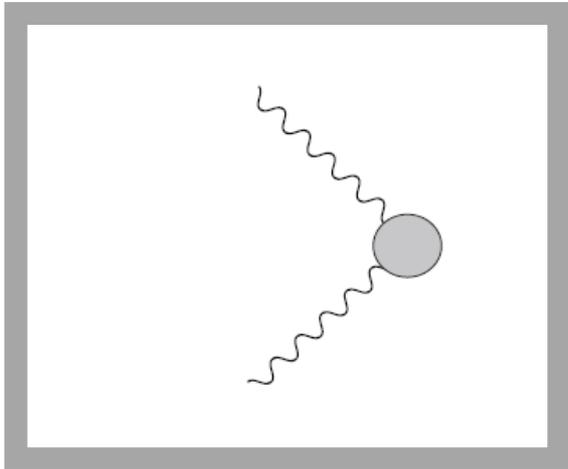
Maximal source magnitude is the critical field

$$\vec{E} \cdot \vec{B} \lesssim (0.2 \text{ T})^2$$

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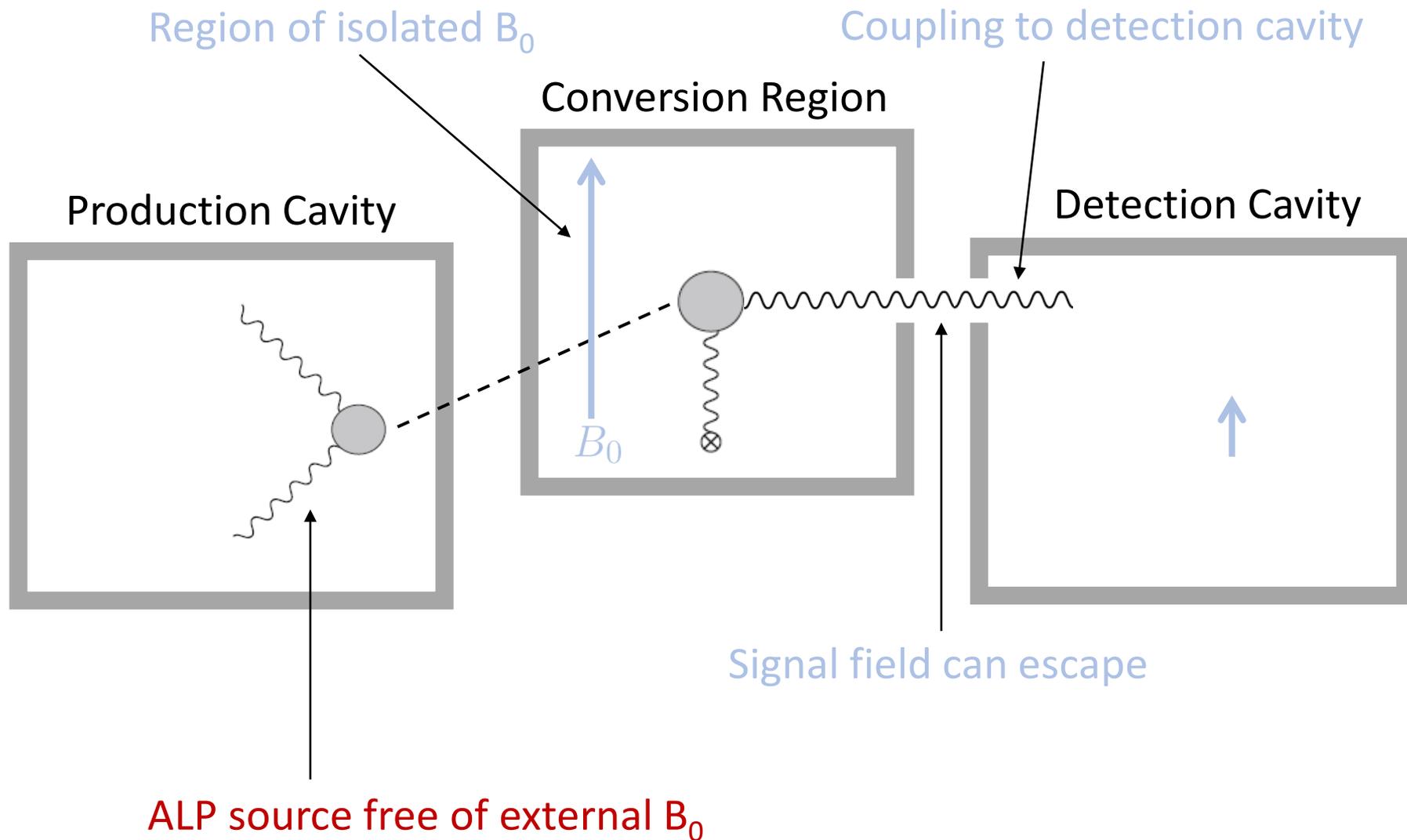
$$\vec{E} \cdot \vec{B} \lesssim (0.2 \text{ T})^2$$

Comparable to conventional cavity with sufficient input power:

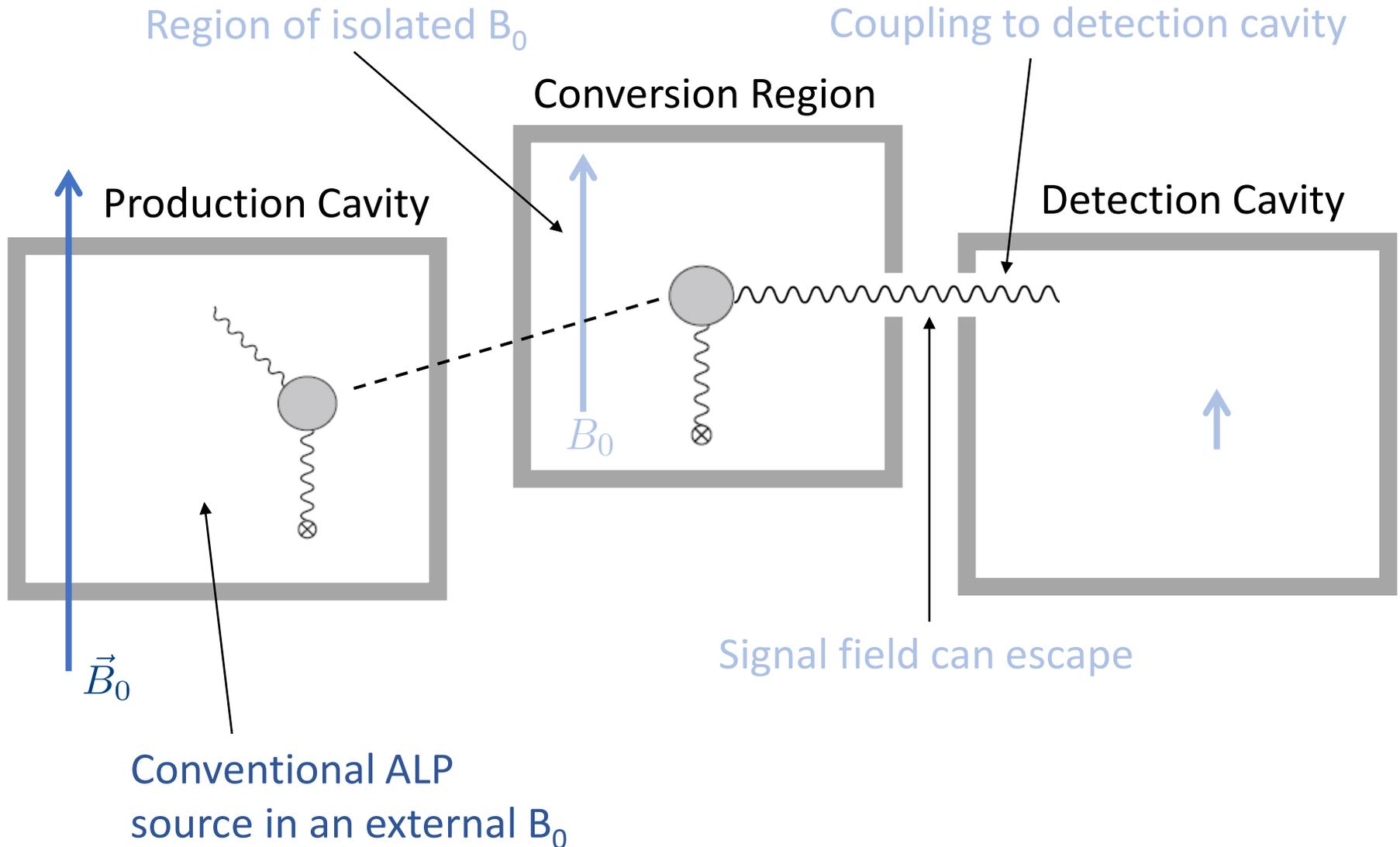
$$(\vec{E} \cdot \vec{B}) \sim (0.1 \text{ T})^2 \left( \frac{P_{\text{input}}}{100 \text{ W}} \right)^{\frac{1}{2}} \left( \frac{Q_{\text{pc}}}{10^5} \right)^{\frac{1}{2}} \left( \frac{B_0}{5 \text{ T}} \right)$$

Fundamental advantage of SRF is in the detection cavity

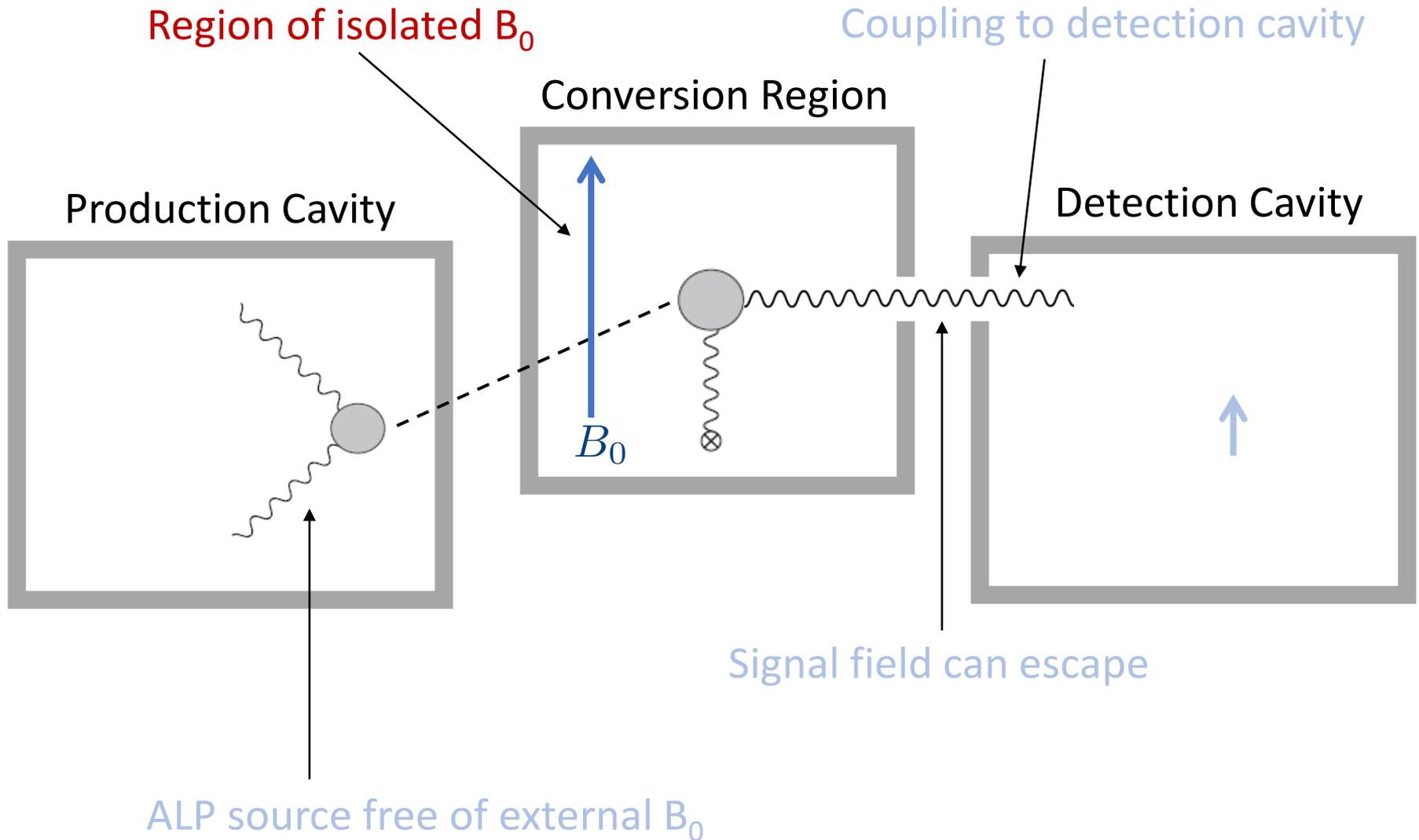
# LSW with SRF Cavities



# LSW with RF and SRF Cavities



# LSW with SRF Cavities



# Gapped Toroid Conversion Region

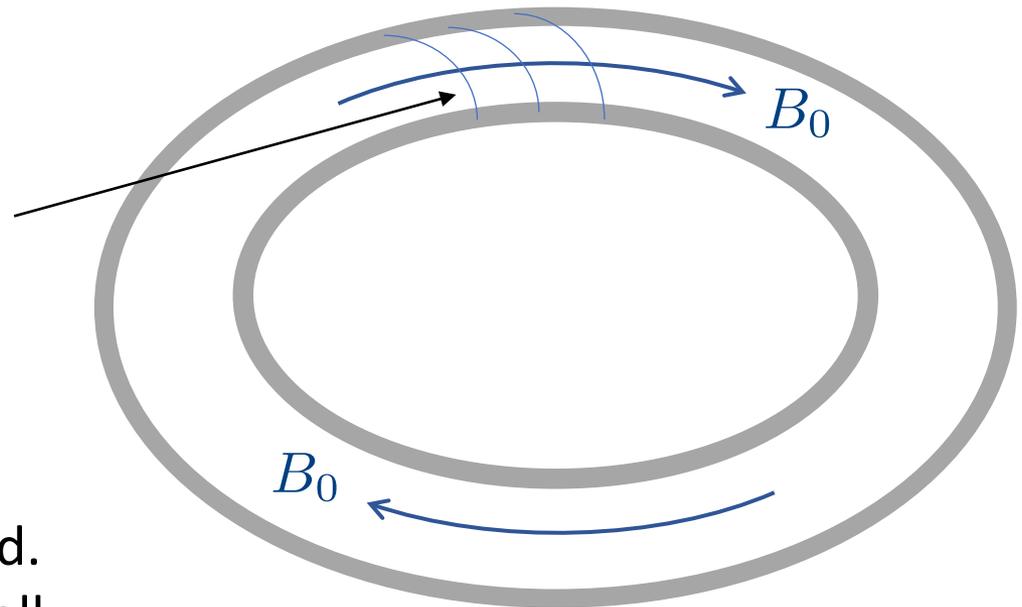
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Confine large static  $B_0$

## Toroidal Magnet

Toroidal frame wrapped with DC current-carrying superconducting wires

Magnetic field (ideally) vanishes outside of toroid.  
In practice, there are small external fringe fields  $\sim 10^{-6} B_0$ .



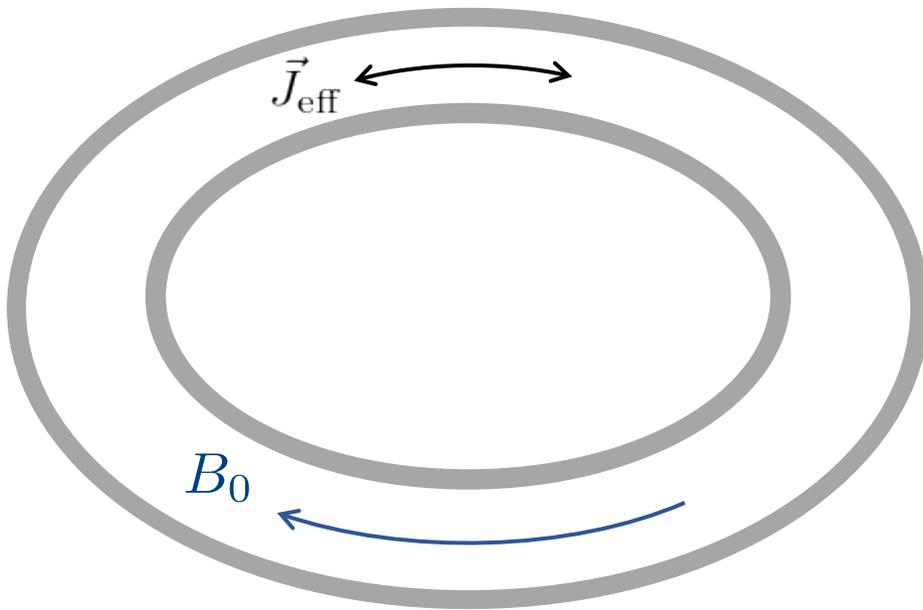
# Gapped Toroid Conversion Region

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## ALP-photon conversion

Axion Electrodynamics:  $\vec{\nabla} \times \vec{B} = \frac{\partial \vec{E}}{\partial t} - g \left( \vec{E} \times \vec{\nabla} a - \vec{B} \frac{\partial a}{\partial t} \right)$

Conversion in  $B_0 \sim$  Effective Current  $\vec{J}_{\text{eff}} = g \vec{B}_0 \frac{\partial a}{\partial t}$

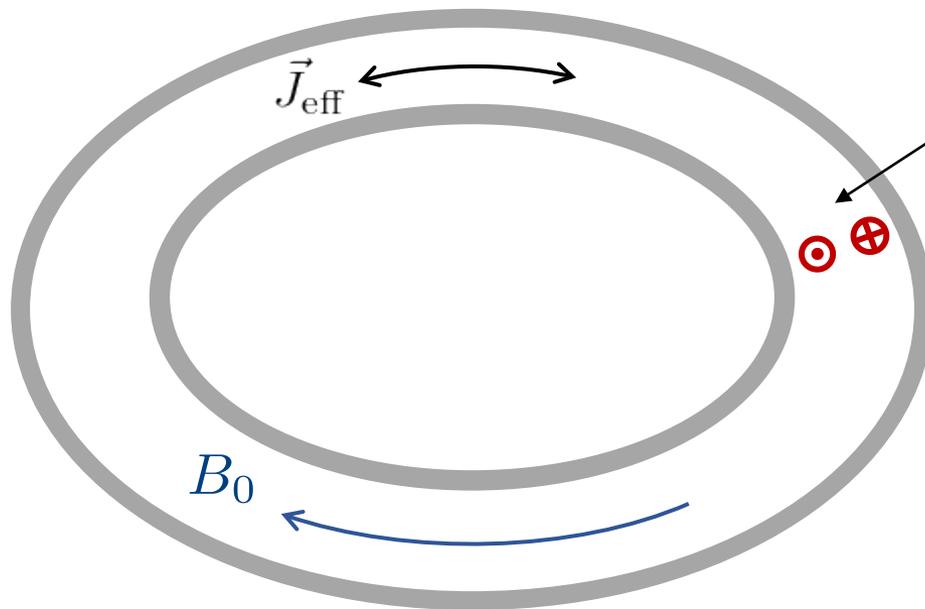


# Gapped Toroid Conversion Region

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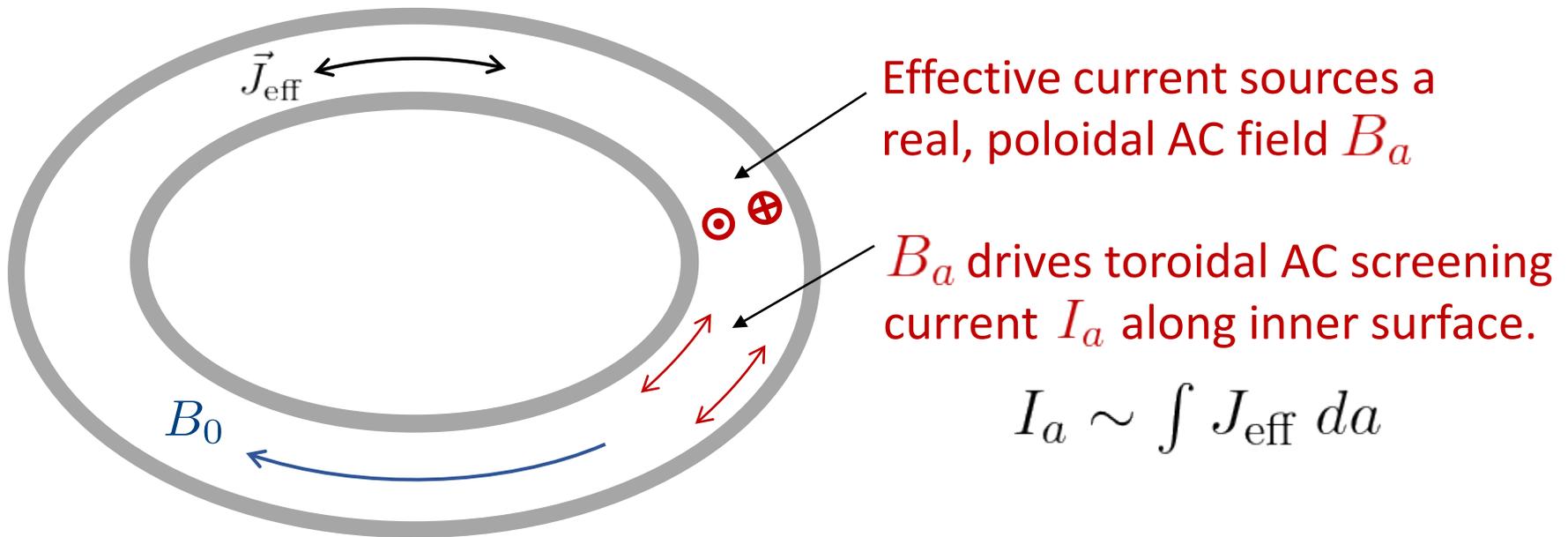
Effective current sources a real, poloidal AC field  $B_a$

# Gapped Toroid Conversion Region

## ALP-photon conversion

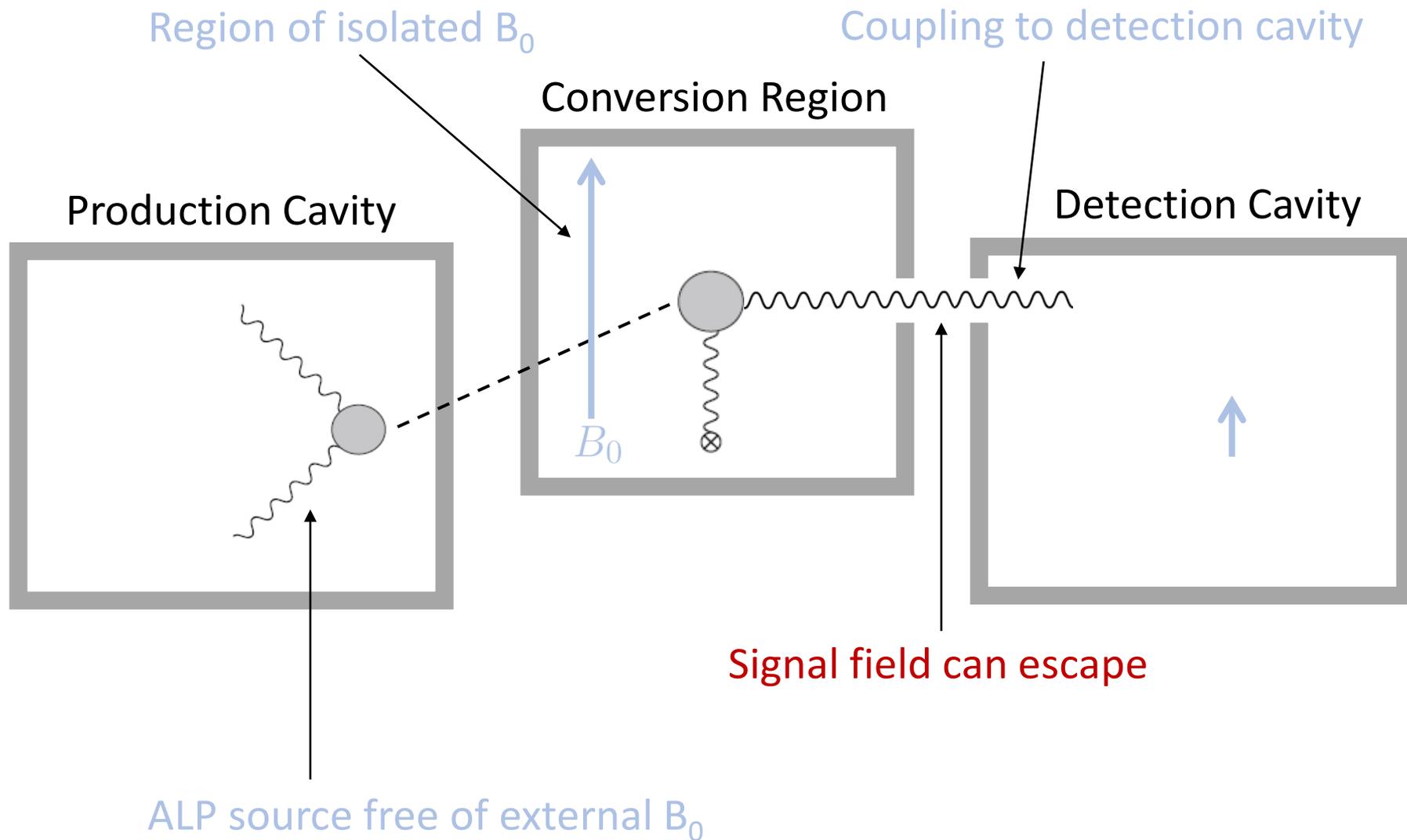
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Conversion in  $B_0 \sim$  Effective Current  $\vec{J}_{\text{eff}} = g \vec{B}_0 \frac{\partial a}{\partial t}$



$$I_a \approx 10^{-13} \text{ nA} \left( \frac{g \text{ GeV}}{10^{-11}} \right)^2 \left( \frac{B_{\text{PC}}}{0.1 \text{ T}} \right)^2 \left( \frac{B_0}{5 \text{ T}} \right) \left( \frac{R}{10 \text{ cm}} \right)^3$$

# LSW with SRF Cavities



# Gapped Toroid Conversion Region

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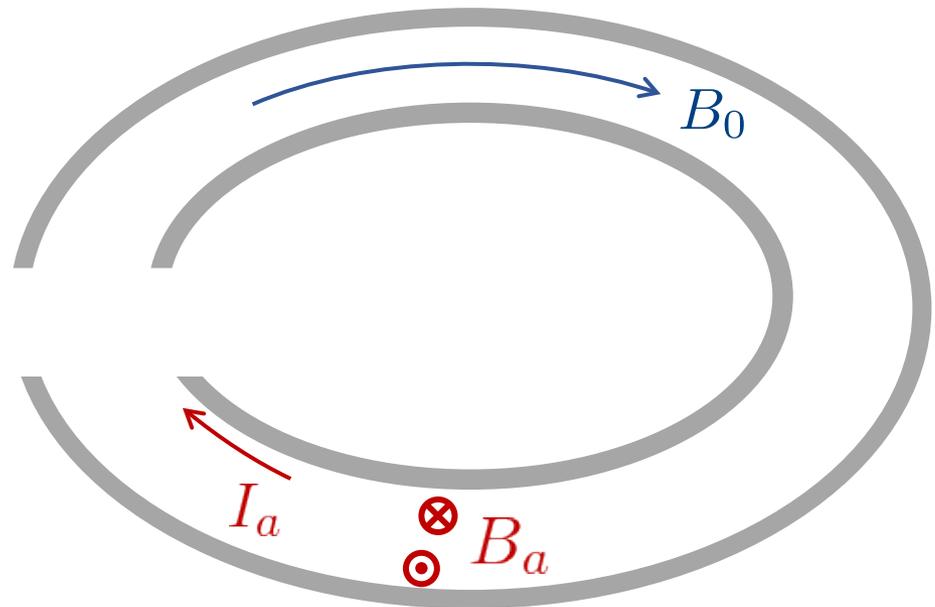
Allow RF signal to escape

Gapped Toroid

Inspired by the dark matter

ALP search ABRACADABRA

[Kahn, Safdi, Thaler '16]



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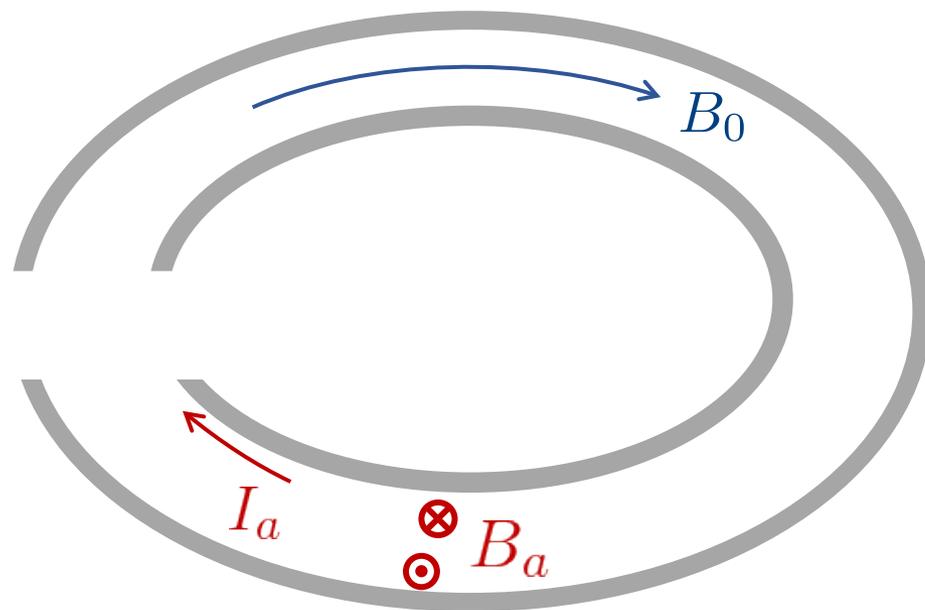
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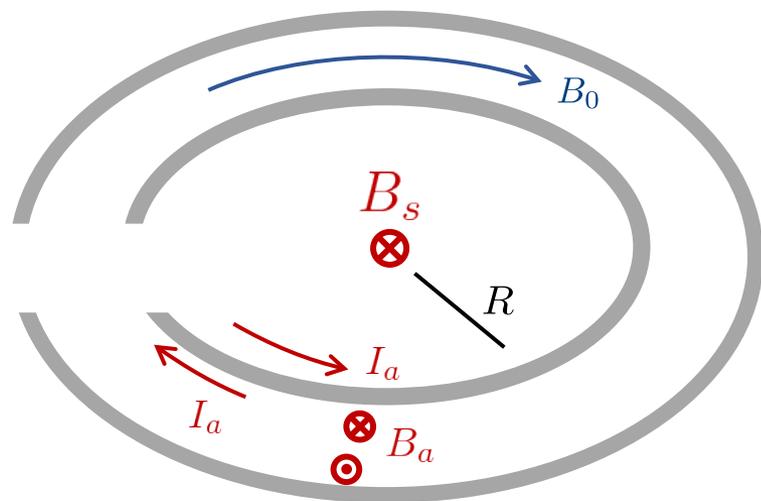
Gap must be large enough to prevent parasitic capacitance, but small enough to minimize external fringe fields.



Propagation of  $I_a$  near the gap depends on the toroid radius  $R$ .

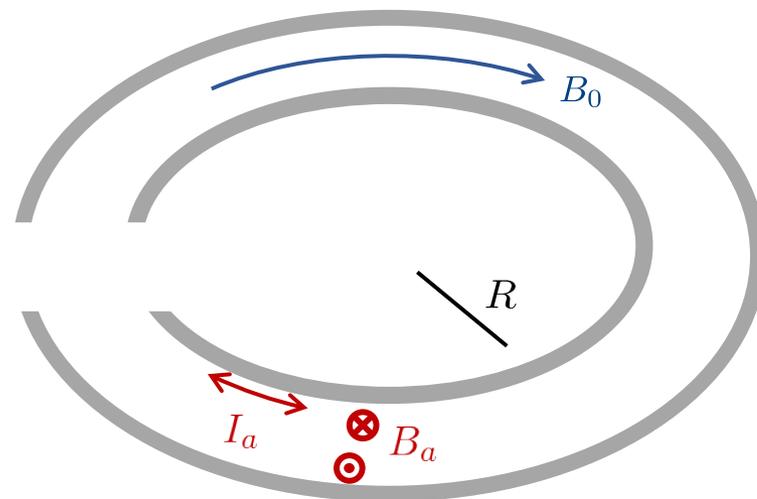
# Gapped Toroid Conversion Region

Allow RF signal to escape



Quasistatic ( $R\omega \ll 1$ )

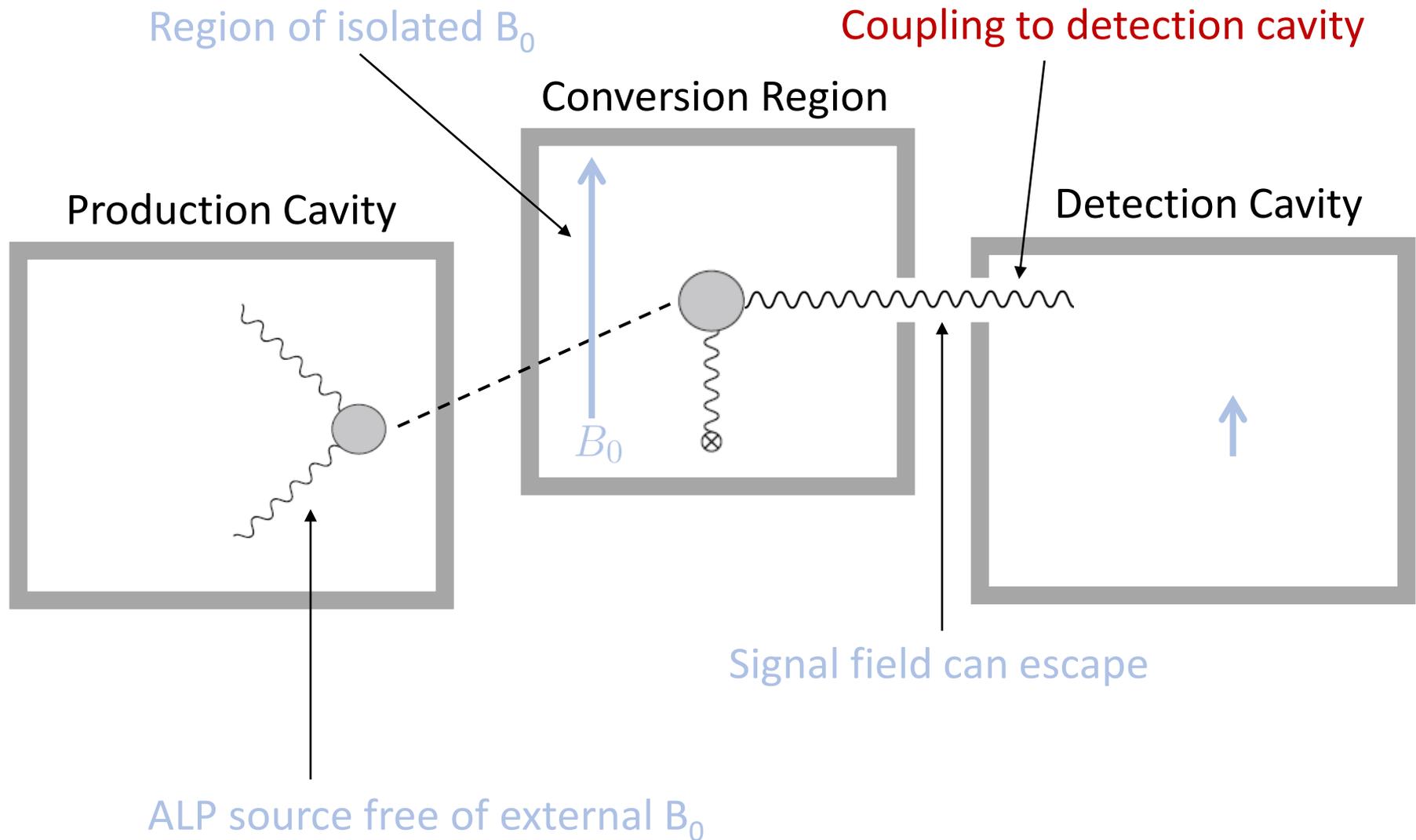
$I_a$  is spatially uniform, returns on outer surface and sources external magnetic field  $B_s$  – “Current Comparator Limit”



Screened ( $R\omega \gg 1$ )

$I_a$  varies rapidly, preventing current from propagating to outer surface and suppressing external field by powers of  $\omega R$ .

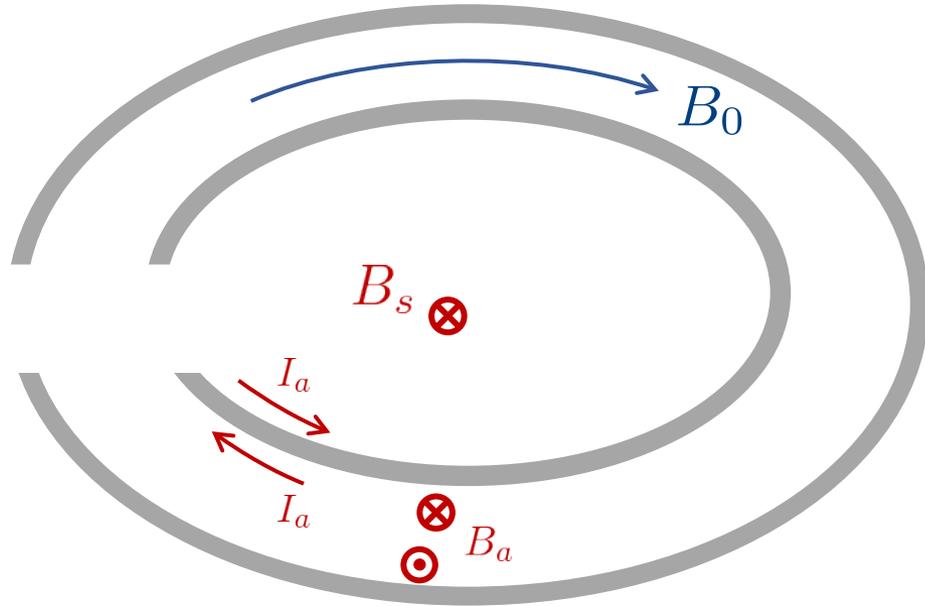
# LSW with SRF Cavities



# Detection Cavity

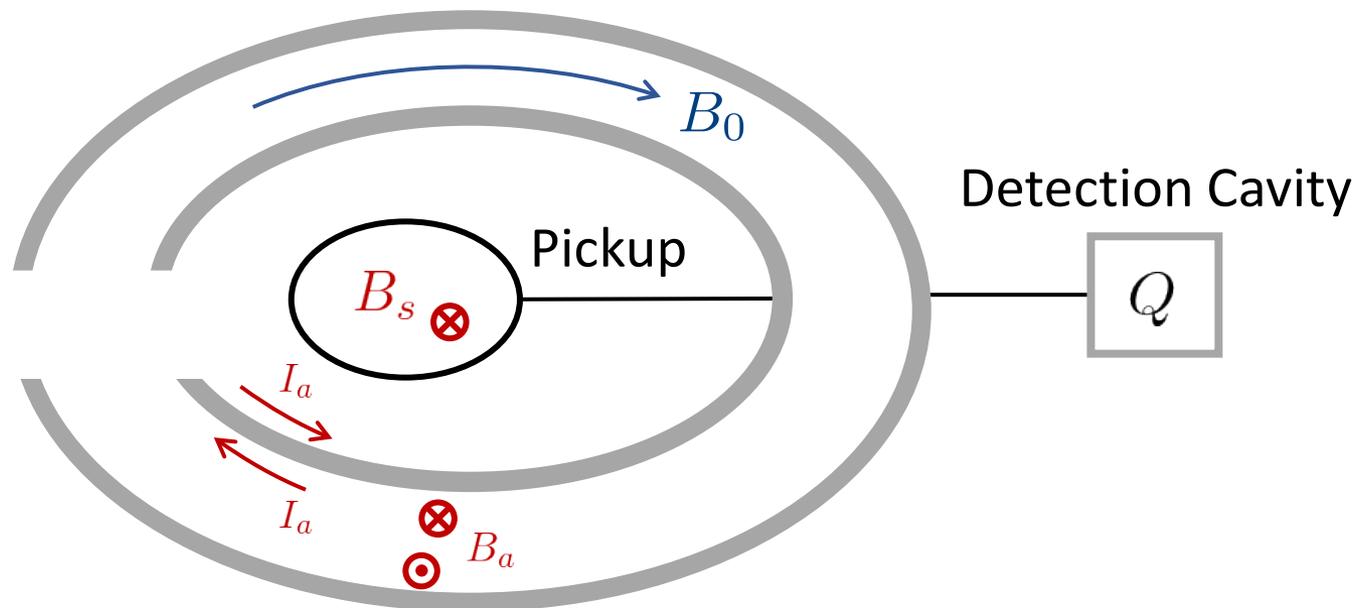
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Couple signal to SRF detection cavity



# Detection Cavity

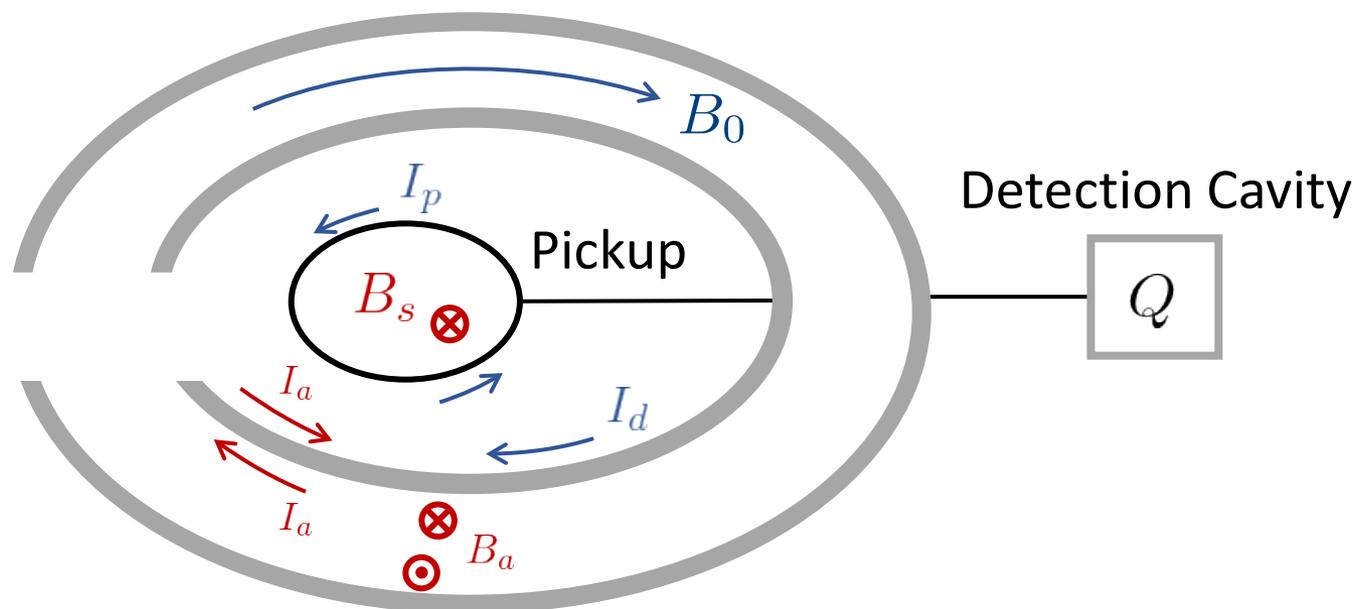
Couple signal to SRF detection cavity



External current  $I_a$  can be coupled to a high-Q SRF detection cavity, e.g. through a pickup loop. A different technique may be preferred in practice, but the optimal signal power is unchanged.

# Detection Cavity

Couple signal to SRF detection cavity



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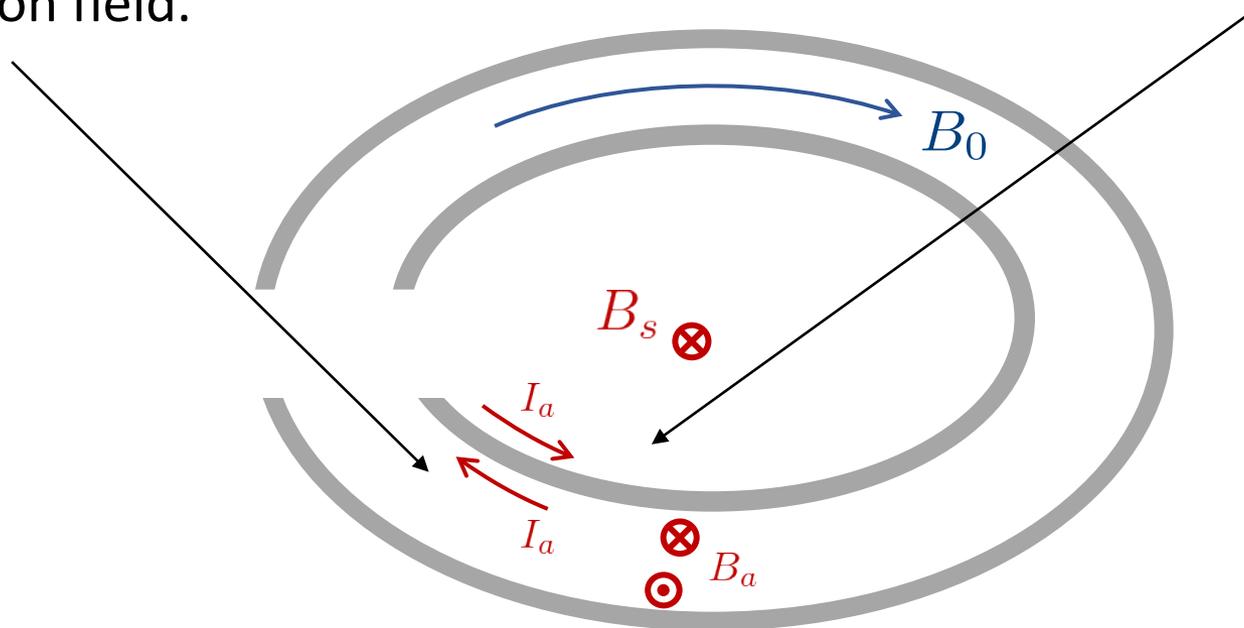
Signal is amplified by  $Q$ , but saturates for sufficiently large  $Q$  due to dissipative backreaction current on the toroid surface.

# Optimal Cavity-Toroid Coupling

How do we maximize the RF signal in the detection cavity?

The inner surface of the toroid is a current source  $I_a$ , supplied by the axion field.

The outer surface of the toroid is an inductor fed by the source  $I_a$ .



Together, the axion field and toroid form a non-ideal current source with a maximum obtainable power draw.

# Optimal Cavity-Toroid Coupling

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Maximum power is extracted when the impedance of the detection cavity matches that of the toroid.

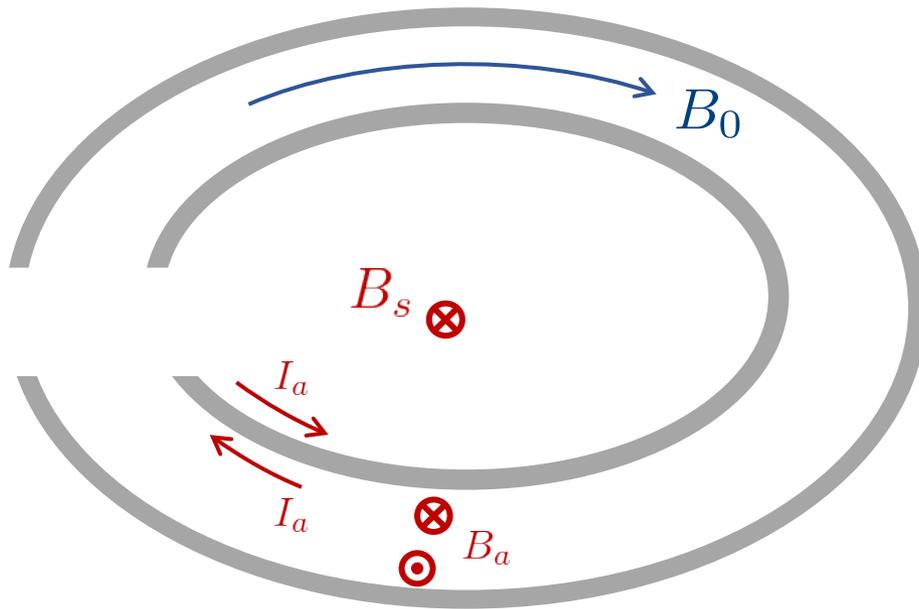
$$\begin{aligned} P_{\max} &= \frac{1}{8} |I_a|^2 \frac{(\omega L_t)^2}{R_t} \begin{array}{l} \longleftarrow \text{Inductance of the toroid} \\ \longleftarrow \text{Resistance of the toroid} \end{array} \\ &\approx 10^{-29} \text{ W} \left( \frac{g \text{ GeV}}{10^{-11}} \right) \left( \frac{100 \text{ n}\Omega}{R_t} \right) \\ &\approx 140 \frac{\text{axions}}{\text{year}} \left( \frac{10^{-5} \text{ eV}}{m_a} \right) \end{aligned}$$

For small Q, impedance matching is not feasible as it demands a pickup inductance which prohibitively perturbs the resonance frequency. We must “undercouple”:

$$P_{\text{uc}} = \frac{1}{2} |I_a|^2 \omega L_t Q \longleftarrow \begin{array}{l} \text{Amplification by} \\ \text{detector cavity Q-factor} \end{array}$$

# Toroid Resistance

Type-II Superconducting wires will contain possibly-dissipative flux tubes due to the large, static conversion magnetic field.



Dissipation is due to the Lorentz force between flux tubes and current:

$$R_t \propto \vec{B}_0 \times \vec{I}$$

Vanishes for an ideal toroid!

Fringe fields will cause  $B_0$  to not be perfectly toroidal, tilted by an angle  $B_{ff}/B_0$ :

$$R_t \approx R_{\perp} \frac{B_{ff}}{B_0} \approx 100 \text{ n}\Omega \left( \frac{B_{ff}}{10^{-6} \text{ T}} \right) \left( \frac{5 \text{ T}}{B_0} \right)$$

# Signal Strength

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Optimal signal power for a given  $Q$  is the minimum of the matched and undercoupled powers

$$P_{\text{signal}} \approx g^4 \frac{B_{\text{PC}}^4 B_0^2}{\omega^6} \text{Min} \left( \frac{\omega L_t}{R_t}, Q \right)$$

Toroid Limited

Cavity Limited

$$\text{Threshold } Q \quad \frac{L_t \omega}{R_t} \approx 10^{10} \left( \frac{R}{10 \text{ cm}} \right) \left( \frac{100 \text{ n}\Omega}{R_t} \right)$$

Superconducting Toroid:  $R_t \gtrsim \text{n}\Omega$

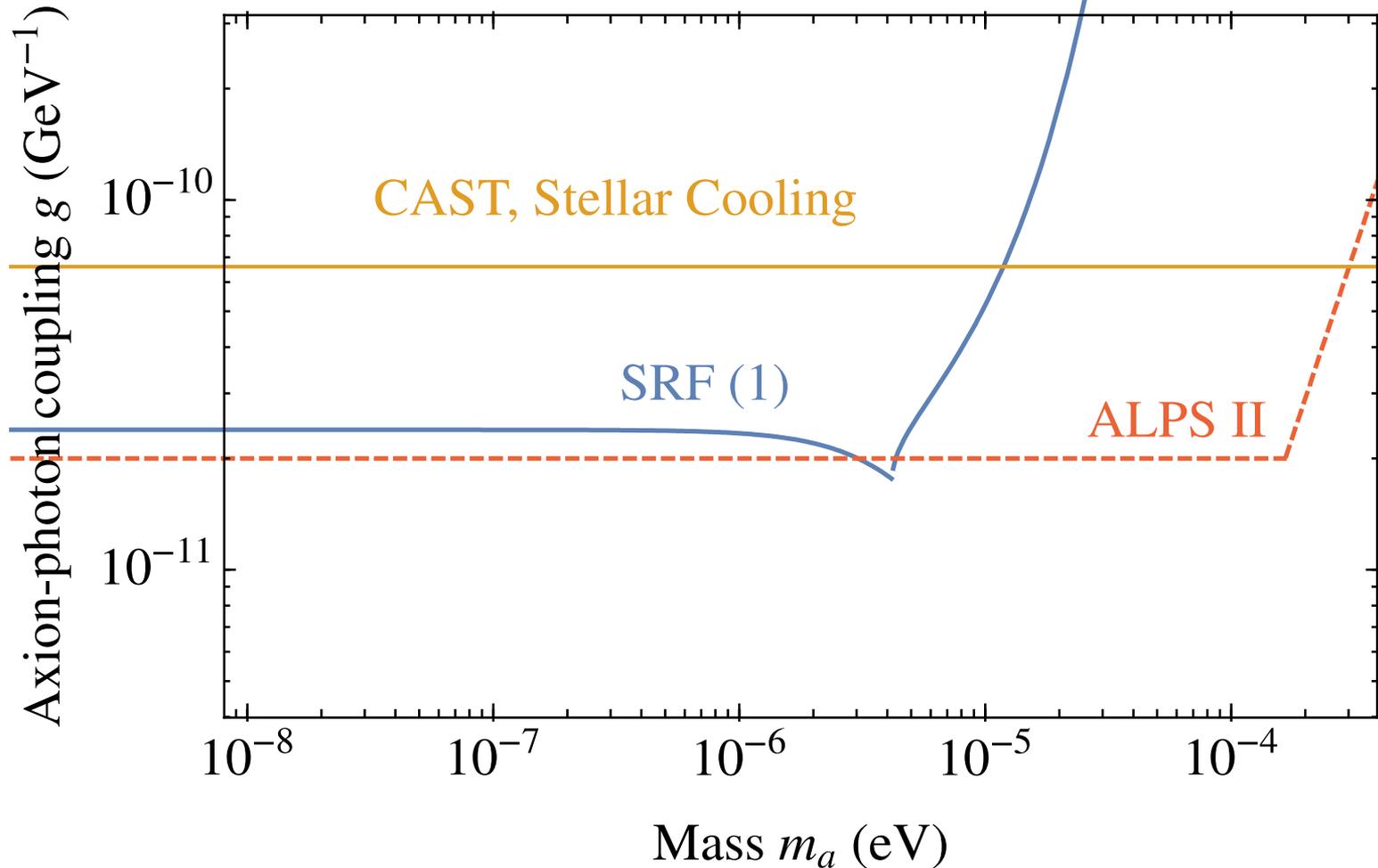
Narrowband noise

$$P_{\text{noise}} = \frac{T_{\text{sys}}}{t_{\text{int}}} \quad T_{\text{sys}} \gtrsim \omega \sim 50 \text{ mK} \quad \text{Quantum Limited}$$

# Projected Sensitivity

$T_{\text{sys}} = 0.1 \text{ K}$ ,  $t_{\text{int}} = 1 \text{ year}$ ,  $R = 10 \text{ cm}$ ,  $\nu = 1 \text{ GHz}$ ,  $B_{\text{PC}} = 0.2 \text{ T}$ ,  $B_0 = 5 \text{ T}$

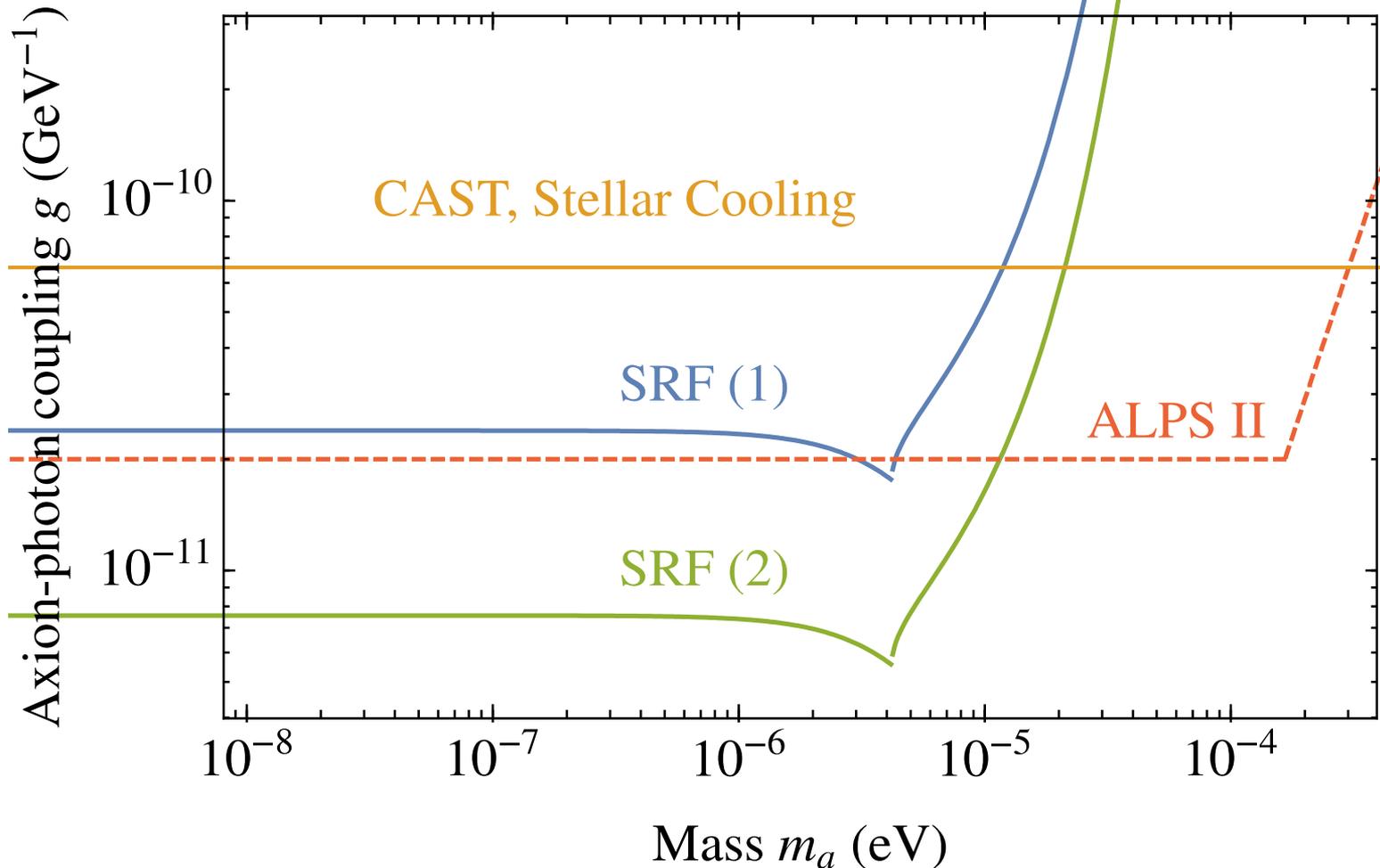
(1)  $R_t = 100 \text{ n}\Omega$  and  $Q \geq 10^{10}$



# Projected Sensitivity

$T_{\text{sys}} = 0.1 \text{ K}$ ,  $t_{\text{int}} = 1 \text{ year}$ ,  $R = 10 \text{ cm}$ ,  $\nu = 1 \text{ GHz}$ ,  $B_{\text{PC}} = 0.2 \text{ T}$ ,  $B_0 = 5 \text{ T}$

(1)  $R_t = 1 \text{ n}\Omega$  and  $Q \geq 10^{12}$



# Axion Production and Detection with Superconducting RF Cavities

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A new design for an LSW ALP search based on SRF cavities

Our realization uses a gapped toroid to confine the magnetic field responsible for ALP-photon conversion, protecting the SRF cavities from quenching.

Major engineering challenges: matching the cavity frequencies and coupling the detection cavity without degrading  $Q$ .

Comparable and complementary to future optical LSW searches and stellar constraints

# Axion Production and Detection with Superconducting RF Cavities

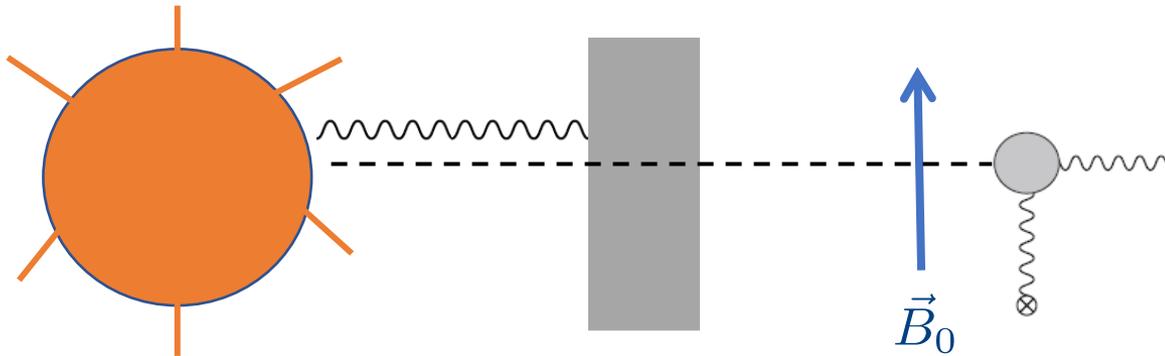
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Extra Slides

# Stellar ALP Searches

“Sun shining through a Wall”

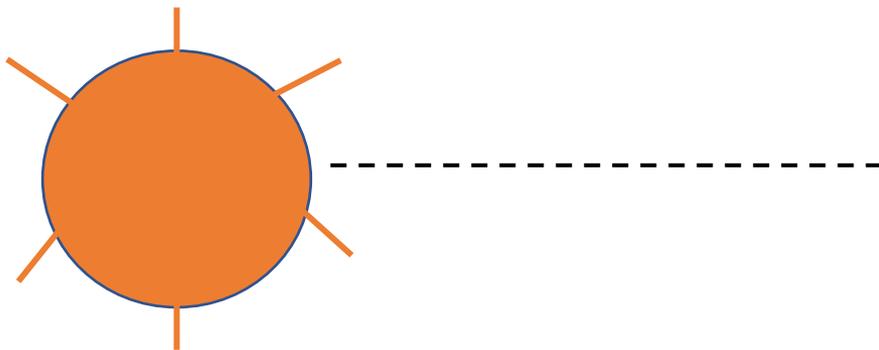
[Sikivie '83, Irastorza et al, '17]



**CAST**

$$g < 7 \times 10^{-11} \text{ GeV}^{-1} \\ \text{for } m_a < \text{eV}$$

Stellar Cooling

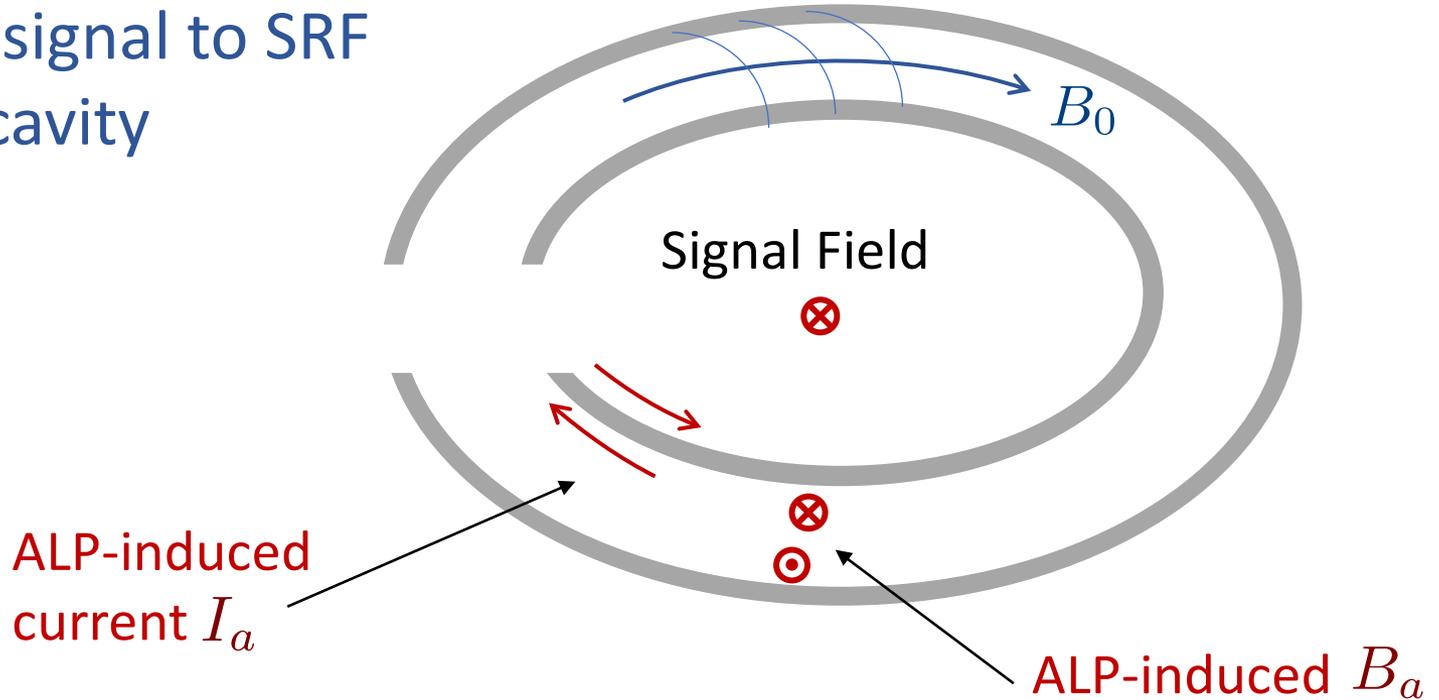


**HB Stars**

$$g < 7 \times 10^{-11} \text{ GeV}^{-1} \\ \text{for } m_a < \text{keV}$$

# Gapped Toroid Conversion Region

(4) Couple signal to SRF detection cavity

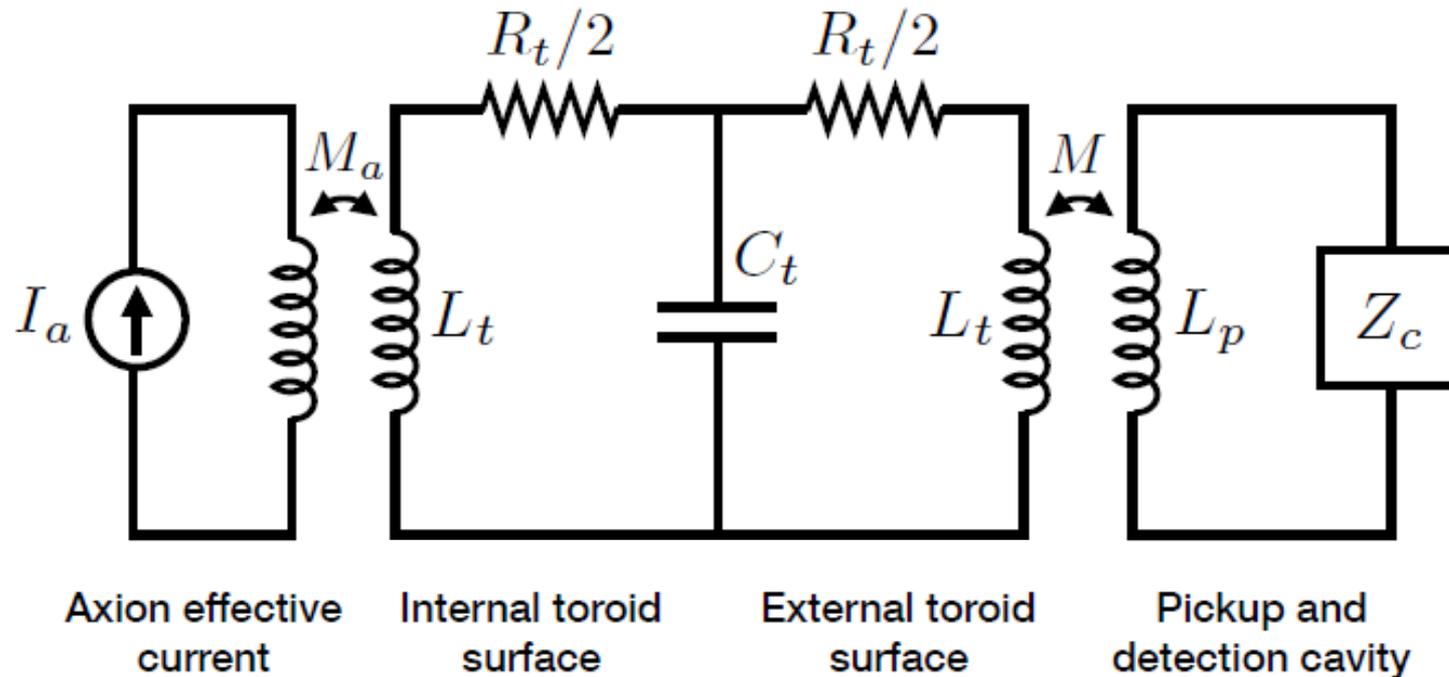


$$I_a \approx 10^{-13} \text{ nA} \left( \frac{g \text{ GeV}}{10^{-11}} \right)^2 \left( \frac{B_{\text{PC}}}{0.1 \text{ T}} \right)^2 \left( \frac{B_0}{5 \text{ T}} \right) \left( \frac{R}{10 \text{ cm}} \right)^3$$

$$B_a \approx 10^{-26} \text{ T} \left( \frac{g \text{ GeV}}{10^{-11}} \right)^2 \left( \frac{B_{\text{PC}}}{0.1 \text{ T}} \right)^2 \left( \frac{B_0}{5 \text{ T}} \right) \left( \frac{R}{10 \text{ cm}} \right)^2$$

# Optimal Signal Power

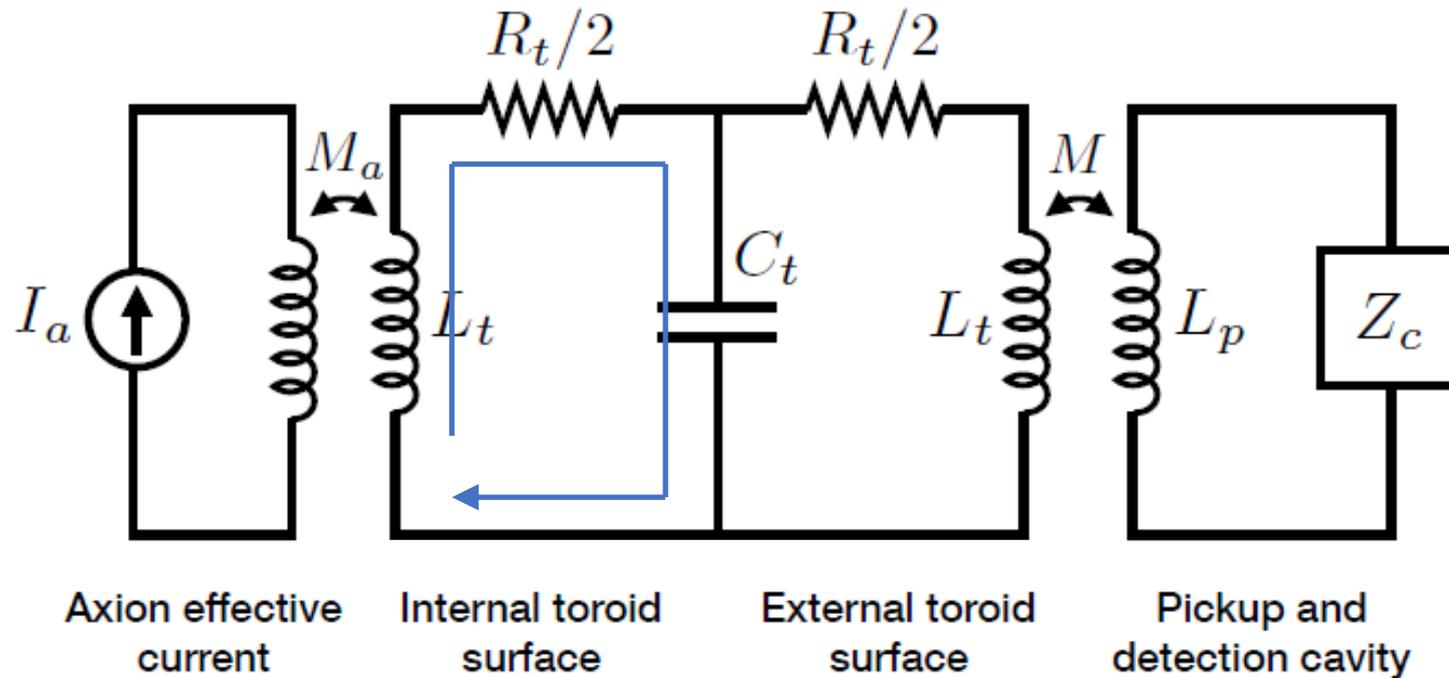
Maximal signal power follows from an equivalent circuit for the axion – toroid – cavity system.



Three current paths on the toroid: completely internal, completely external, or wrapping internal and external

# Optimal Signal Power

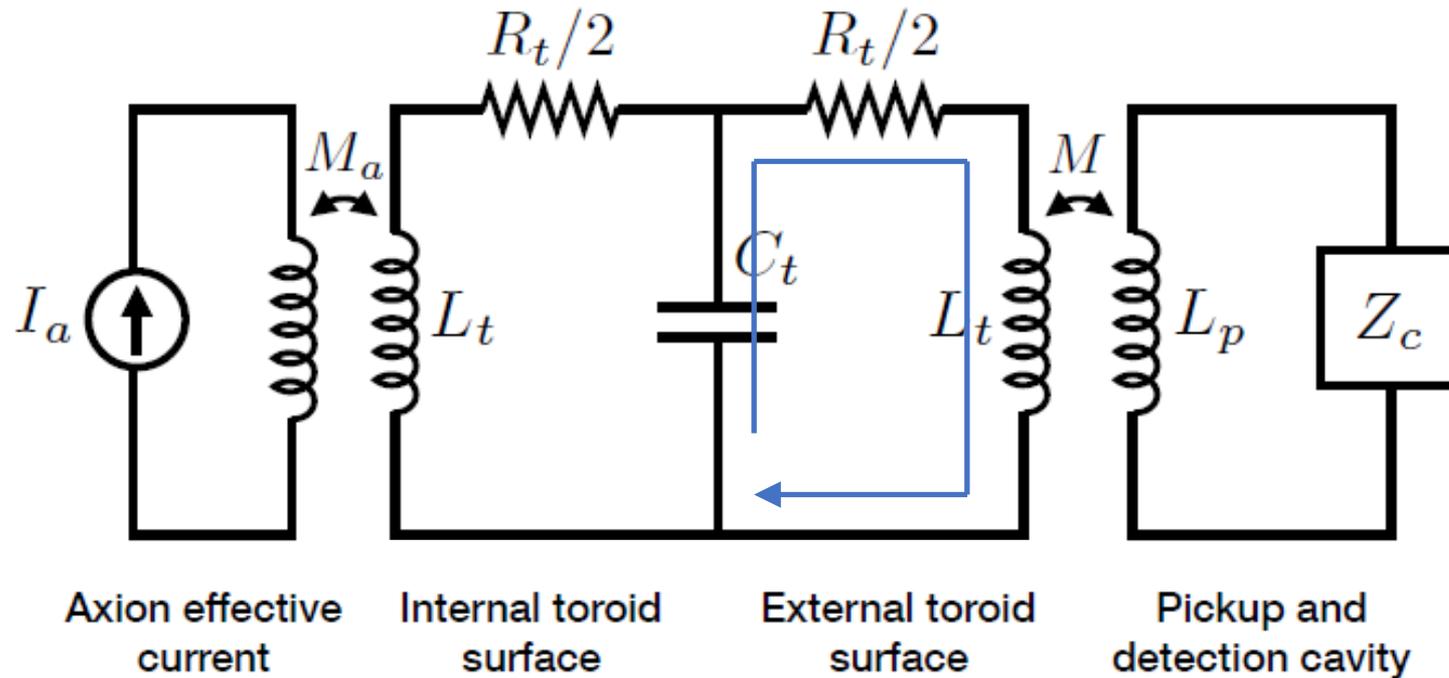
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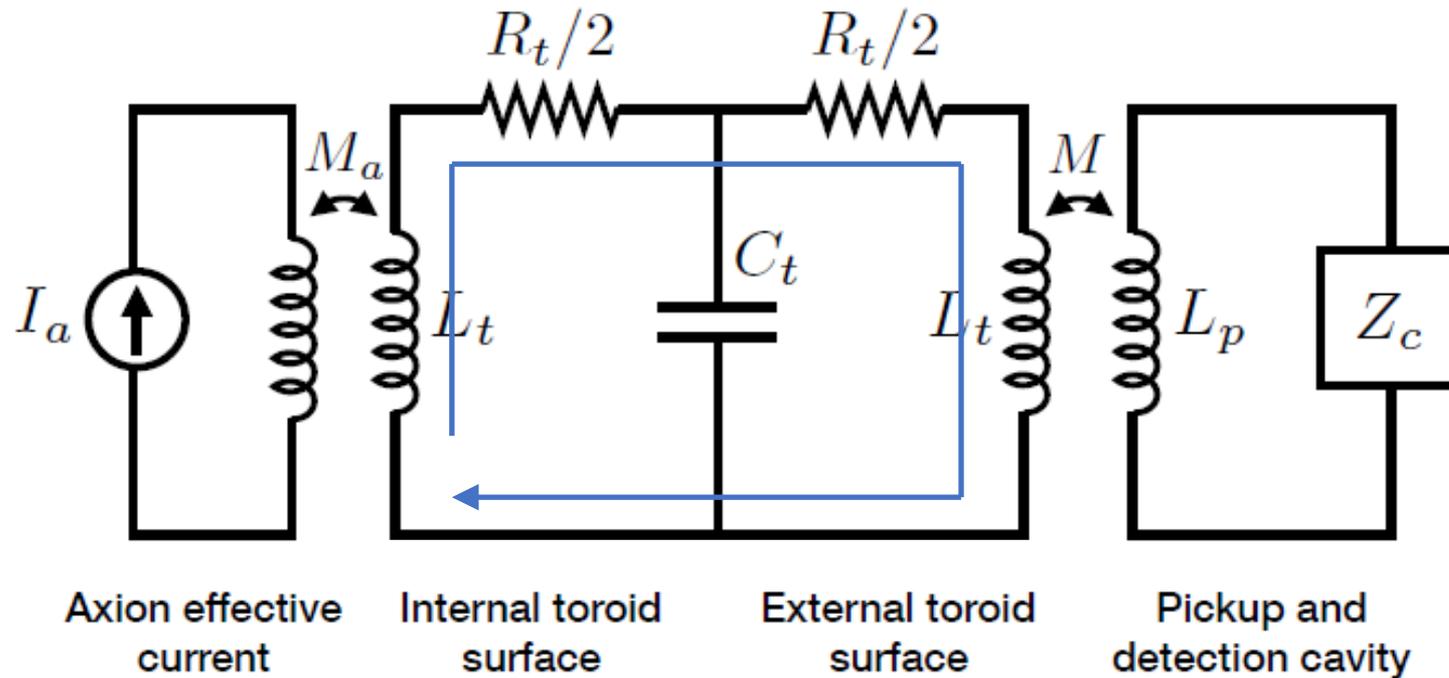
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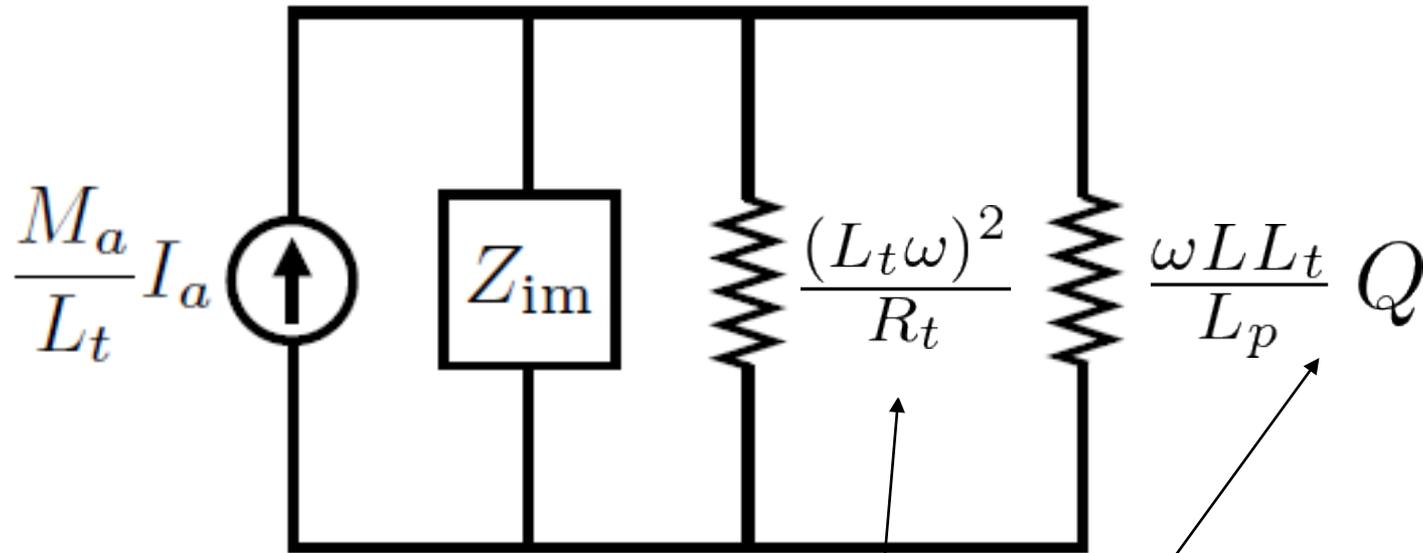
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# Optimal Signal Power



Circuit simplifies:

Resistor for dissipation in the toroid

Resistor for dissipation detection cavity

Imaginary impedance (determines resonance frequency)

Maximum power is transferred to the cavity when  $L_p$  is tuned such that the two resistance are equal.

# Optimal Signal Power

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Tuning  $L_p$  alters the resonance frequency, which is given by the zero of  $Z_{im}$

$$Z_{im} \sim \left( \frac{2}{i\omega L_t} + i\omega C_t + \frac{1}{i\omega \frac{L_t}{L_p} L} + i\omega \frac{L_p}{L_t} C \right)^{-1}$$

$$\omega_{res} \sim \omega_0 \sqrt{1 + 2 \frac{L}{L_p}}$$

Require  $L_p \gg L$  to preserve resonance near GHz – otherwise it is not the detection cavity that will ring up, but the pickup loop.

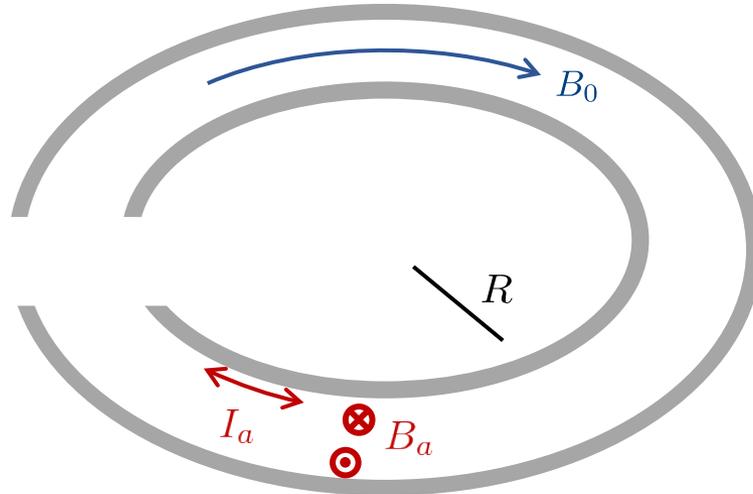
This obstructs impedance matching for small  $Q$ , as the optimal  $L_p$  is

$$L_p^{\text{match}} = \frac{L}{L_t \omega} R_t Q \gg L$$

$$Q \gg \frac{L_t \omega}{R_t}$$

# AC Screening

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Increasing the frequency above the quasistatic limit will suppress the external current and external field.

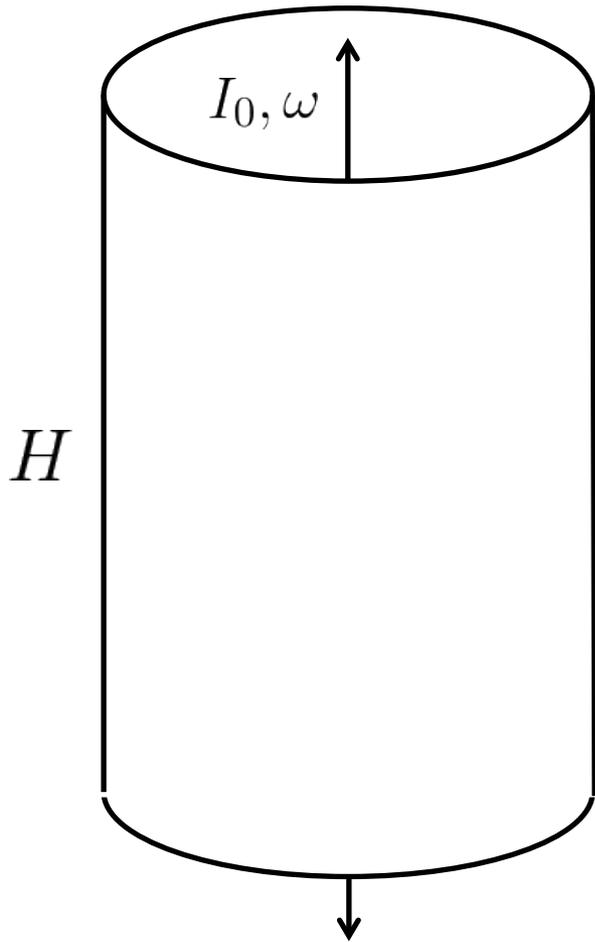
Suppression will scale as a power-law in frequency.

This is hard to calculate in the full toroid geometry, but we can see the nature of the effect in a simpler case.

# Toy Model of Screening

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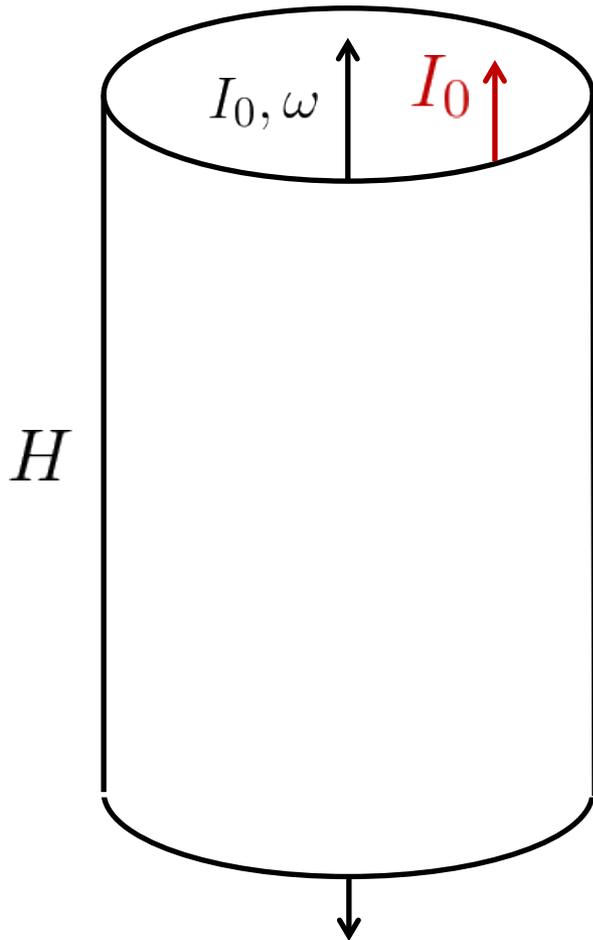
An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



# Toy Model of Screening

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An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?

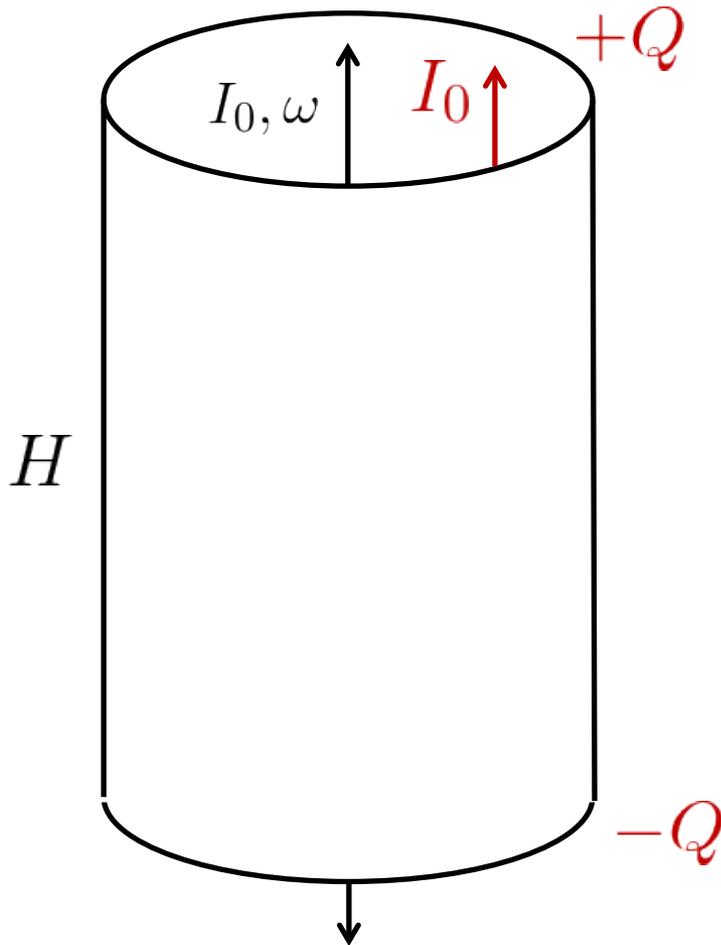


A screening current must first be established on the inner surface.

# Toy Model of Screening

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An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?

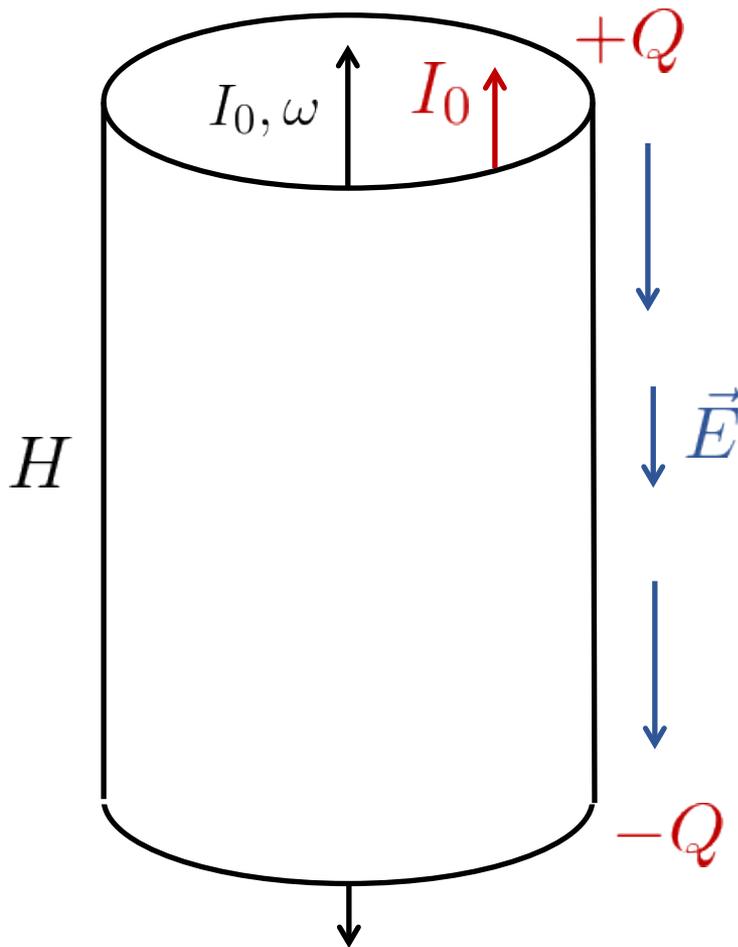


A screening current must first be established on the inner surface.

This causes a build-up of charge on the edges, which backreacts on the screening current

# Toy Model of Screening

An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



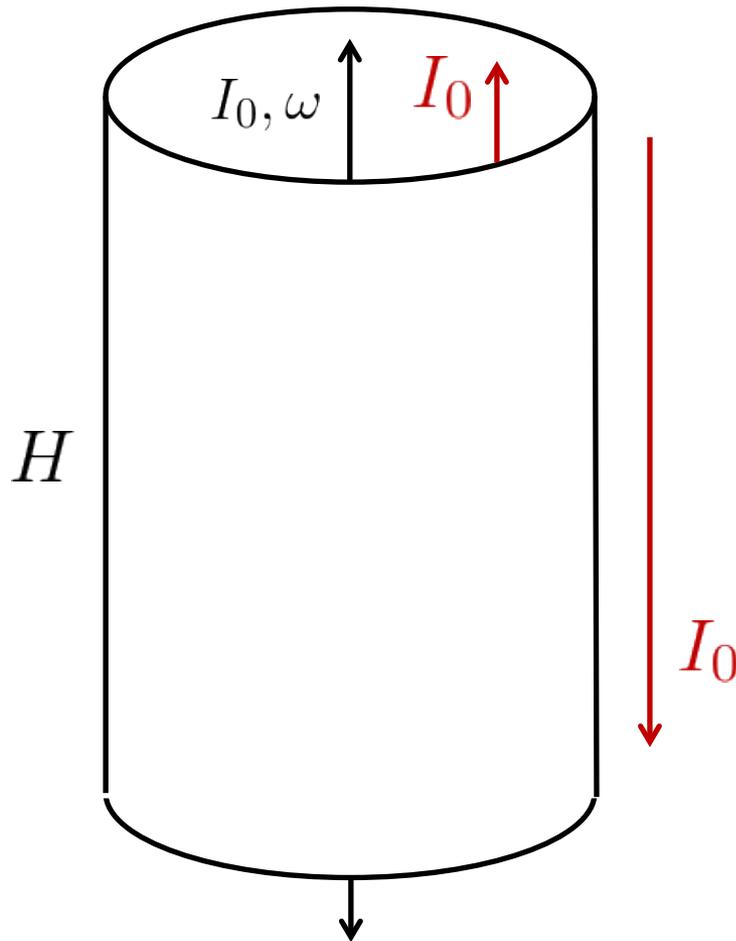
Quasistatic ( $\omega H \ll 1$ )

Edge charges source quasistatic  
Coulomb field, with uniform  
direction along the cylinder.

# Toy Model of Screening

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An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



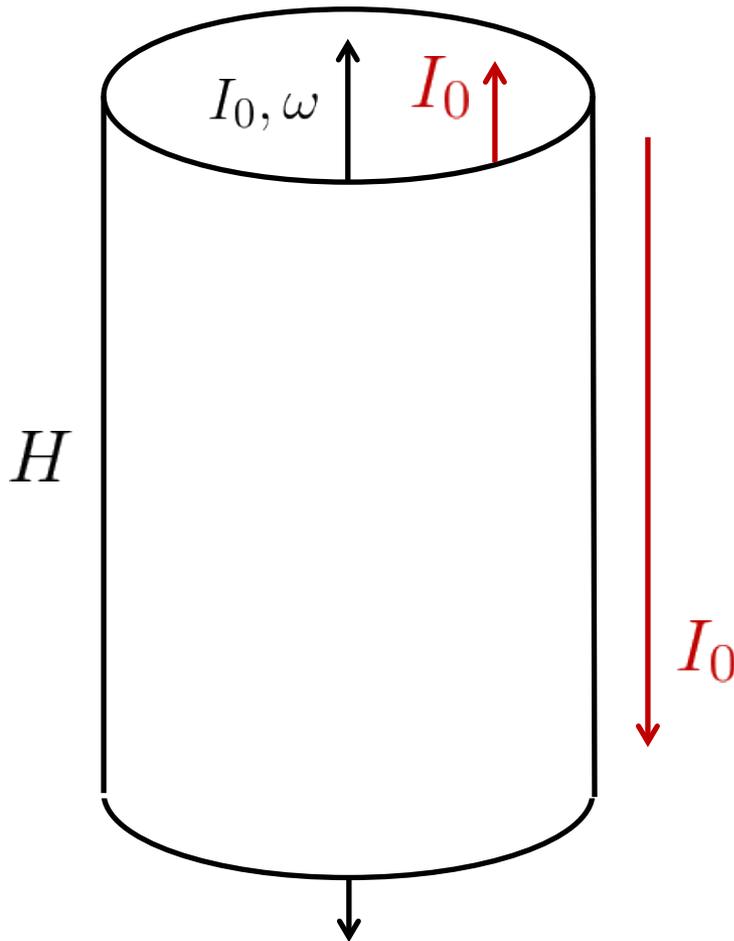
Quasistatic ( $\omega H \ll 1$ )

Edge charges source quasistatic Coulomb field, with uniform direction along the cylinder.

Coulomb field drives uniform external current, eliminating charge build-up.

# Toy Model of Screening

An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



Quasistatic ( $\omega H \ll 1$ )

Edge charges source quasistatic Coulomb field, with uniform direction along the cylinder.

Coulomb field drives uniform external current, eliminating charge build-up.

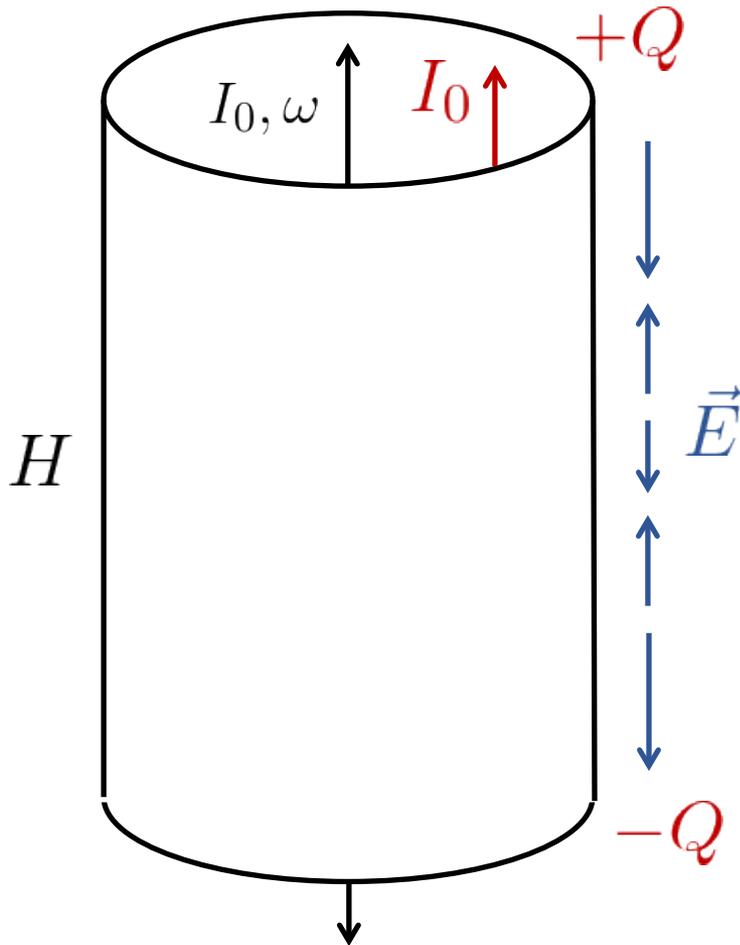
External current sources a field equal to that of the central current.

No Screening

# Toy Model of Screening

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An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?

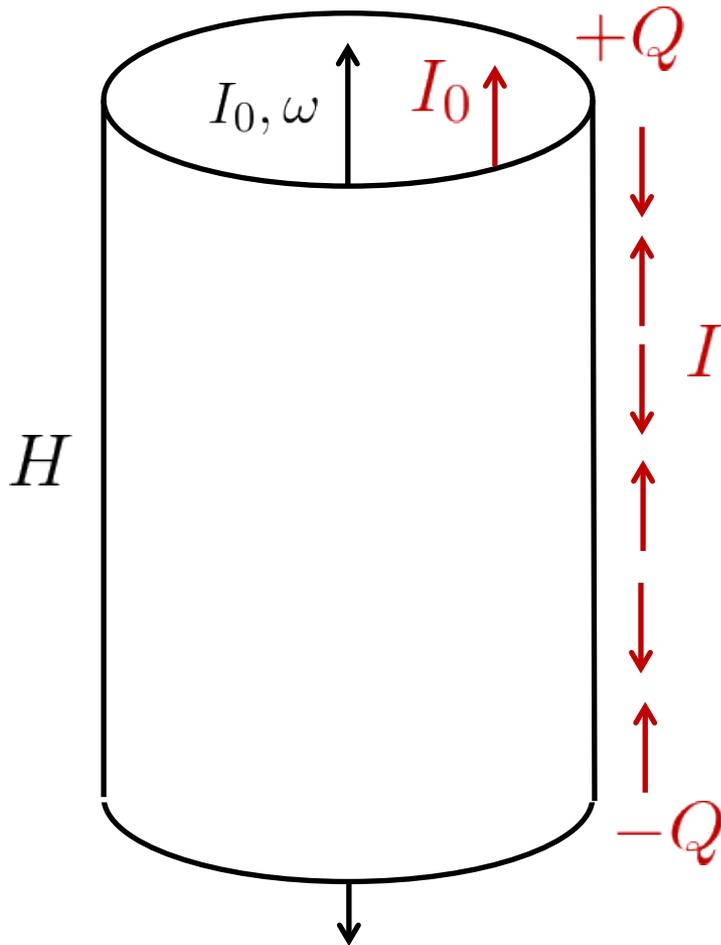


Screened ( $\omega H \gg 1$ )

Edge charges source radiative electric field, with alternating direction.

# Toy Model of Screening

An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



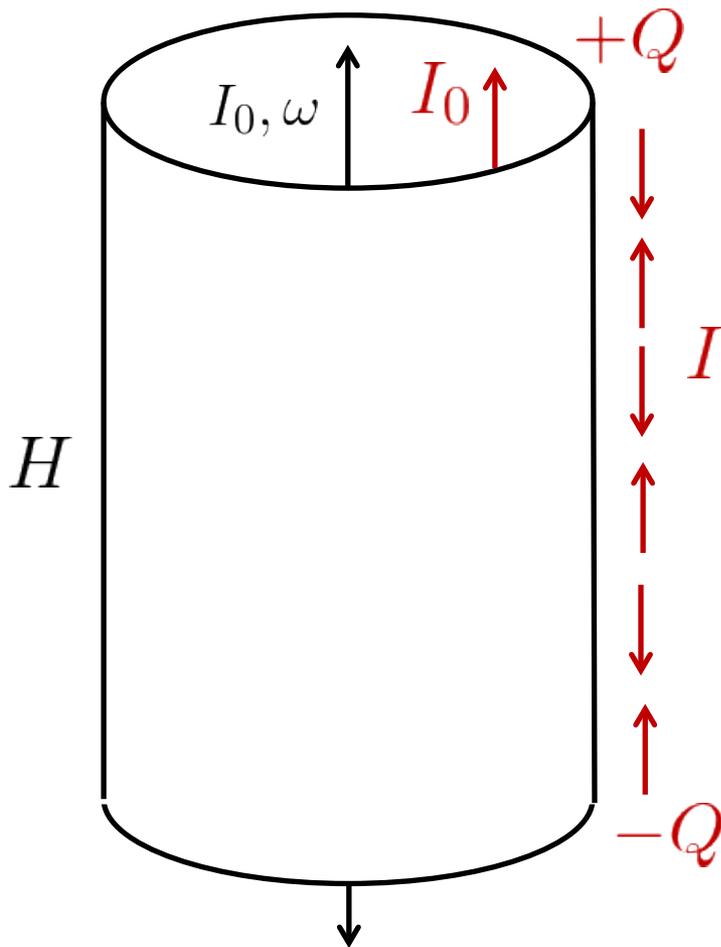
Screened ( $\omega H \gg 1$ )

Edge charges source radiative electric field, with alternating direction.

Radiative field drives spatially alternating current, charge build-up is maintained.

# Toy Model of Screening

An AC current is switched on inside of a finite cylindrical, conducting shell – what are the external fields in the vicinity of the conductor?



Screened ( $\omega H \gg 1$ )

Edge charges source radiative electric field, with alternating direction.

Radiative field drives spatially alternating current, charge build-up is maintained.

Non-uniform external current sources a multi-pole field, suppressed relative to the field of the internal current.

Power-law Screening